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Resonant holographic measurements of laser ablation plume expansion in vacuum and argon gas backgrounds

Lindley, Roger Alan, Ph.D.

The University of Michigan, 1993

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RESONANT HOLOGRAPHIC MEASUREMENTS OF LASER ABLATION PLUME EXPANSION IN VACUUM AND ARGON GAS BACKGROUNDS

by

Roger Alan Lindley

A dissertation submitted in partial fulfillment of the requirements for the degree of Doctor of Philosophy (Nuclear Engineering) in The University of Michigan 1993

Doctoral Committee:

Professor Ronald Gilgenbach, Chair Associate Professor Mary Brake Professor Terry Kammash Professor Y.Y. Lau Professor Duncan Steel ©<u>Roger Alan Lindley</u> 1993 All Rights Reserved This dissertation is dedicated to my Master, Jesus Christ, of whom I am greatly indebted. It has been by His leading that I have studied at the University of Michigan and performed this work.

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CHAPTER 1 INTRODUCTION

Introduction to Laser Ablation

Laser ablation is the removal of material from a solid surface by an energetic laser beam. Figure 1.1 shows a schematic of the basic process. An energetic laser beam is focused onto a solid surface where some of the energy is absorbed. Surface material is ejected away from the surface on the order of 10 km/s due to mechanisms dependent upon the target



Fig. 1.1 Schematic of Laser Ablation and Laser Ablation Deposition. T=Target, AP=Ablation Plume, TF=Deposited Thin Film, and S=Substrate.

material. A shallow crater then remains on the target surface where the laser beam was incident. If there is a background gas of reasonable pressure (≥ 200 mTorr), the expanding vapor is inhibited by collisions with the background gas, causing condensation of the vaporized material as well as fluid dynamic vortex formations (e.g., a mushroom cloud shape) (Sap92). At lower pressures, the ablated material can then be collected on another surface by placing a substrate near the target. Generally, the substrate is heated to > 50 °C to improve the material qualities of the deposited thin films.

The physical mechanisms removing material from the target surface vary with the type of target used. For the ablation of metals, the dominant mechanism is thermal heating, where the temperature of the target surface increases dramatically causing liquefication and vaporization (Dre91). A second ablation process for metals, observed under certain conditions, is called photochemical ablation (Dre91). In this process, the laser photons are absorbed by optically opaque electronic sites in the polycrystal structure of the metal. This energy then directly breaks the bonds of some of the atoms in the crystal, expelling them from the surface.

While the ablation of metals is dominated by the thermal ablation mechanism, both thermal ablation and photochemical ablation are significant mechanisms in the laser ablation of polymers (i.e., plastics). Certain sites on polymer chains are optically opaque and thus provide absorption sites in the lattice at which to break molecular bonds (Sri93, Dre86). For example, Srinivasan observed a considerable yield of MMA (methylmethacrylate) polymer chains in the ablation plume following the ablation of PMMA (polymethylmethacrylate) (Sri93). Both mechanisms were used to explain the observed plume composition.

For the ablation of ceramic materials, however, photochemical ablation has been shown to be the dominant mechanism (Dre86, Dic90). Defects in the inorganic lattice are optically opaque, providing locations for the laser energy to break lattice bonds, and it has been shown that an increase in the surface defect concentration increases the ablative material expulsion for the same laser ablation conditions (Dic90).

The physical mechanisms removing material from the target surface depend not only upon the target material but also on the laser fluence, the laser pulselength, and the laser wavelength. If the laser fluence is high enough and/or the laser pulselength is long enough, the ablation plume forms a hot, dense plasma, causing the laser energy to be absorbed by the plasma electrons via inverse bremsstrahlung. The absorbed laser energy is then transmitted to the target by heat conduction. This third ablation process is called plasma etching, although it is not truly separate from the thermal ablation process (Dre91).

Finally, the laser wavelength makes a difference in the ablation process. Ultraviolet (UV) lasers have photon energies (4.9 eV for 248 nm) that are on the order of molecular bond energies (e.g., ≈ 1.9 eV for Al-Al and ≈ 3.6 eV for C-C), making photochemical ablation processes more efficient. Visible and infrared (IR) lasers have lower energies per photon and thus require multiple-photon-absorption for the photochemical ablation mechanism. This in turn means that more of the laser energy is thermally absorbed by the target, that the effectiveness of photochemical ablation is reduced, and that the effectiveness of thermal ablation is enhanced (Bra91, Con89). However, there are two main drawbacks of using visible and especially IR lasers (Bra91). First of all, the rapidly, thermally heated region of the target is not confined to the laser spot. This causes unwanted surface damage (cratering) to the target when attempting to ablate desired patterns in the target surface (laser machining). Secondly, large debris is ejected from the surface instead of small fragments or single atoms. Not only is the large debris undesirable for making thin films on substrates, it also has a tendency to land on the target just beyond the laser spot. Again, this is undesirable for laser machining, especially for electronics. Therefore, UV lasers are preferred for both laser machining and laser ablation deposition.

Applications of Laser Ablation

As indicated above, the first application of laser ablation is the micromachining of metals, plastics, and ceramics, and the second application of laser ablation is the deposition of thin films onto substrates. Laser ablation is being intensely studied and used in the computer chip industry for drilling printed circuit boards (IBM), as a method of etching circuit patterns in photoresist (lithography) in preparation for plasma etching, etching circuit patterns in the semiconductor materials as a replacement for RF-plasma etching, and depositing semiconductor and superconductor materials onto silicon wafers. Micromachining is also useful in surgery, most notably for kidney stones (lithotripsy) and for eyes (corneal and keratorefractive surgery).

Laser ablation with very intense laser energies is being investigated as a possible X-ray source for X-ray lithography, as a possible commercial source of electricity by inducing fusion reactions (Inertial Confinement Fusion), and as a possible laser-gain medium for X-ray lasers. The production of X-ray lasers is especially appealing to biologists since the electromagnetic absorption by proteins at a wavelength around 4 nm is much greater than for water (Elt90), thus allowing holographic observation of living cells and tissue. This would certainly revolutionize biological investigations that presently rely upon diagnostics requiring non-living samples. Laser ablation with very low laser intensities is already being utilized for nondestructive materials testing and the labeling/marking of plastic, metal, ceramic and glass surfaces via laser-induced surface effects.

Table 1.1(a) and (b) show a selection of materials versus diagnostics for which laser ablation plume expansion has been investigated and Table 1.2 shows a selection of targets and substrates that have been studied using laser ablation deposition. The diagnostics in Table 1.1 will be discussed in the next section.

The targets have been chosen either because they have a strong resonant transition at a wavelength accessible by the then-current-day dye laser technology or because of their present or possible future use in making electronic circuits. Barium (Ba) and potassium (K) are examples of the former. As examples of the latter, aluminum (Al), copper (Cu), silicon (Si), silicon compounds, metallic oxides, metallic fluorides, and YBCO (YBa₂Cu₃O_{7-x}, a superconducting ceramic) are of interest in making thin films. In addition, the ablation of plastics (such as lucite, mylar, PET, PMMA, polyimide and polystyrene) is studied to develop better spatial control over the removal of printed circuit board insulating layers and photoresists in the process of making of computer chips.

	Al	Ba	Cu	Si	graphite	AlO	BN	YBCO	CaF LiF MgO	NaO SiO	GaAs	Polyi- mide	PET	рмма	lucite	poly- sty- rene
Ch.Part. Collector					Mea88					Dic90						
Langmuir Probe			Gut89													
TOF Mass Spectro.					Cro91			Geo89			Wan92					Fel88
Laser Defl.					Enl87 Mea88				Pet88			Ven92b	Sel91 Ven90	Ven90	Ven92b	
Abs. Spectr.			Sap91					Che91 Geo89								
Hook Spectr.			Sap93													
Em. Spectr.	Rot84b			Rot84a	Bra87 Mea88	Rot84a Rot84b	Geo93	Geo93 Zhe89								
LIF	Wan91		Dre91			Dre86		Oka92								
PLIF		Cap90	Sap92													

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Table 1.1(a) Diagnostics versus ablation target materials studied; a selection of articles investigating laser plume expansion.

	Al	Ba	Cu	Na	Ge MgO Si	BN	K Li/Pb	YBCO	mylar	PET	PMMA	polyi- mide	AlO
R.&T. Scatt.		Mos88	Sap91										
Em. Photo.						Geo 93		Geo92 Geo93					
TOFPRL				Bak87									
Shadow -graphy								Gup91			Sri93		
Absorption Photography	Gil91 Ven 92a	Jel81 Mos88 Sie77							Ven92a			Ven92a	
Schlieren	Ven92a								Ven92a	Ven90	Ven90	Ven92a	
Resonant Schlieren	Jel81 Koo79 Mos88 Sie77												
Michelson Interfer- ometry	Wal86		Wal86		Wal86						Wal86	Wal86	Wal86
Resonant Holography	Lin93	K0079					Dre75						

Table 1.1(b) Diagnostics versus ablation target materials studied, continued; a selection of articles investigating laser plume expansion.

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Ref.	Target	Substrate	Ref.	Target	Substrate
Col 92	Diamond	Ti	Niu 92	SrCuO	SrTiO
Eri	AlN,Co	Si	Nor 91	AIN	Sapphire
F = 00	Dy	C:	Qia 93	(cBN)	Si
Fog 90	SIO	51	Raj 93	CoLZnTe	Ga-compounds
Kar 92	YBCO	SrTiO	Roy 92	BaTiO	Pt
K0191	CeO SrtiO	Si and SrTiO	Sat 88	Graphite	Si, Quartz
Kre 93	Cu,Fe and Co alloys, PdSi	Si and AlO	Sch 69	Cr, W, C Ti, ZnS SrTio SbSZnS	Glass
	Sn		Tar 93	InO	In-compounds MgO, YSZ
Kwo 88	YBCO	SrTiO, Zr MgO, Sapphire	Ven 88	увсо	Si
Liu 92	BaCaCuO	MgO	Zhu 93	BiSrCa- CuO BiSrCuO	MgO

Table 1.2 Target materials versus substrate materials; a selection of articles investigating laser ablation deposition.

The Study of Plume Expansion

This dissertation focuses on taking resonant holograms of expanding plumes from the laser ablation of solid aluminum targets and the subsequent calculation of line-integrated-densities of the neutral aluminum atoms in the ground state within the ablation plumes from those holograms. Since many different diagnostics have been used to study laser ablation plume expansion, a summary of the diagnostics utilized and the parameters that each measure are presented here.

Table 1.3(a) and (b) show various diagnostics versus the parameters they measure for a selection of articles investigating the expansion of laser ablation plumes. The simplest diagnostic is Charged Particle Collection, where the ablation particle current from a DCvoltage biased collection plate is measured as a function of time and bias voltage (Dic90, Mea88). This gives the total number of charged particles collected, here called the ion yield, as well as the time of flight (TOF) most-probable expansion velocity of the variously charged plume components, as selected by the bias voltage.

When the collection surface is instead the tip of a very small wire, charged particle collection is a Langmuir probe measurement (Gut89). Again, the collected current is measured as a function of DC-bias voltage and time, but here Langmuir probe theory can be used to calculate the particle density and the electron temperature of the ablation plume. Although the total ion yield is no longer measurable, the current as a function of time still gives charged particle TOF, most-probable expansion velocities.

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	species recog.	fluid dyn.	atomic dens.	dens. grad.	expan. vel.	kinetic temp., f(v)	inter. temp.	elec. temp.	part. yield	thresh. fluence	mach #	thermal diff.
Chg.Part. Collector	Dic90				Dic90 Mea88				Dic90 Mea88	, , , , , , , , , , , , , , , , , , ,		
Lang.probe	Gut89		Gut89		Gut89			Gut89				
TOF Mass Spectro.	Cro91 Fel87				Cro91 Fel87	Cro91 Fel87			Cro91 Fel87 Wan92			
Laser Defl.	Enl87 Mea88		Enl87 Ven90	Enl87 Mea88 Pet88 Sel91 Ven90 Ven92b	En187 Mea88				Enl87 Mea88	Mea88 Pet88 Sel91		Sel91
Abs. Spectro.	Che91 Geo89 Sap91		Che91 Sap91		Che91 Geo89	Geo89						
Hook Spectro.	Sap93		Sap93			_						
Em. Spectr.	Bra87 Geo93 Mea88 Rot84b Zhe89				Bra87 Geo89 Mea88 Zhe89	Zhe89	Geo93	Bra87 Mea88		Rot84b	Zhe89	
LIF	Dre86 Dre91 Sap91 Wan91				Dre86 Dre91 Sap91 Wan91	Dre86 Dre91 Wan91	Dre86 Sap91 Wan91					
PLIF	Cap90 Sap92	Cap90 Sap92	Cap90 Sap92	Cap90	Cap90 Sap92							

Table 1.3(a) Diagnostics versus parameters measured, a selection of articles investigating laser ablation plume expansion.

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	species recog.	fluid dyn.	atomic dens.	elec. dens.	expan. vel.	kinetic temp. f(v)	init. target temp.
R. & T. Scatt.	Mos88 Sap91		Mos88		Sap91		
Em. Photo.	Geo92 Geo93	Geo92 Geo93			Geo92 Geo93		
TOFPRL					Bak87	Bak87	
Shadow- graphy		Gup91 Sri93			Gup91 Sri93		Gup91
Abs. Photo.	Gil91 Jel81 Mos88 Sie77 Ven92a	Gi191 Je181 Mos88 Sie77 Ven92a			Gil91 Je181 Mos88 Sie77 Ven92a		
Schl.		Ven90 Ven92a			Ven90 Ven92a		
Res. Schl.	Jel81 Koo79 Mos88 Sie77	Jel81 Koo79 Mos88 Sie77			Jel81 Koo79 Mos88 Sie77		
Mich. Int.		Wal86	Wal86	Wal86	Wal86		
Res. Holo.	Dre75 Koo79 Lin 93	Dre75 Koo79 Lin93	Dre75 Koo79 Lin93	Dre75 Koo79	Dre75 Koo79 Lin93		

Table 1.3(b) Diagnostics versus parameters measured, continued; a selection of articles investigating laser ablation plume expansion.

A TOF Mass Spectrometer placed in the path of the expanding ablation plume measures the TOF particle flux as a function of mass and time (Cro91, Fel87, Wan92). This allows the TOF most-probable expansion velocity, the velocity distribution, and thus the kinetic temperature to be determined as a function of mass.

If an unexpanded laser beam is passed through an ablation plume with strong density gradients, the laser beam will be deflected (Laser Deflection) (Enl87, Mea88, Pet88, Sel91, Ven90, Ven92b). The amount of deflection can be used to measure the path-integrated density-gradient of the ablation plume. When cylindrical symmetry is assumed for the ablation plume, the particle density can also be calculated (Enl87).

When the wavelength of the unexpanded, laser beam is tuned near an atomic or molecular resonance and passed through an ablation plume with weak density gradients, the transmission as a function of wavelength can be measured (Absorption Spectroscopy) (Che91, Geo89, Sap91). The line-width and line-shape of the transmission signal can provide information concerning the line-broadening collisional mechanisms occurring in the ablation plume. This allows the atomic density to be calculated from the transmission signal at peak absorption. Density versus time then gives the TOF, most-probable expansion velocity, the velocity distribution, and thus the kinetic temperature of the plume.

If the unexpanded, dye laser beam passing through the ablation plume (with weak density gradients) has a wide bandwidth, the relative phase shift of the laser light as a function of wavelength can be measured (Hook Spectroscopy) (Sap93). A peak in the phase shift occurs on either side of a resonance line, so the resonant line-density along the beam path can be calculated from the wavelength separation of the two peak phase shifts.

During the ablation process, atoms and molecules are given kinetic energy while molecules are given rotational and vibrational energy (internal energy) as well. In addition, electronic states within the atoms are excited and some electrons are removed from their atoms. As the electrons and ions recombine and the excited electronic states relax, observable light is emitted. Emission Spectroscopy collects and measures the light emitted from a single point in space as a function of wavelength (Bra87, Geo93, Mea88, Rot84b, Zhe89). Light emitted at known atomic and molecular resonance wavelengths provides species-recognition and molecular internal-temperatures. The intensity of light at the peak of a resonance as a function of time provides the TOF, most-probable expansion velocity, the velocity distribution, the fraction of the ionic species and thus the kinetic temperature of the plume.

When a dye laser is focused onto the same point in space described above and the wavelength is tuned to the resonance of a plume constituent, an upward electronic transition can be induced in a large fraction of the ablated atoms and the intensity at a known relaxation wavelength can be observed versus time. This is the principle of Laser Induced Fluorescence (LIF), and it normally utilizes certain transitions in given atoms and ions where a strong relaxation signal can be observed at a wavelength other than the induced excitation wavelength (Dre86, Dre91, Sap91, Wan91). The measured emission signal versus time gives the TOF most probable velocity, the velocity distribution, and thus the kinetic temperature of the plume. In addition, emission spectroscopy can be done simultaneously at other wavelengths to obtain molecular-internaltemperatures and the velocity measurements for other atomic constituents.

If a dye laser beam tuned to an LIF excitation wavelength is focused into a sheet of light (a plane) and passed through the plume, a photograph of the LIF-relaxation light emitted from this plume can be taken perpendicular to the plane of light (Planar Laser Induced Fluorescence, PLIF) (Cap90, Sap92). This provides expansion fluid dynamics and leading edge expansion velocities for the resonantly excited plume constituent.

When the dye laser light passes through the ablation plume, there is a small portion of photons that are radiatively scattered by the electrons (Thomson Scattering), elastically scattered by the atoms and molecules (Rayleigh Scattering) (Mos88), and inelastically scattered by the atoms and molecules (Raman Scattering). Sappey and Gamble (Sap91) have measured Thompson and Rayleigh scattering signals on top of their LIF signal and utilized them to determine the location and velocity of the electrons and molecules, respectfully, in their ablation plumes.

When a camera is situated to view the expansion of the ablation plume away from the target, the plume image can be captured on film or electronically stored using a CCD camera (Emission Photography) (Geo92, Geo93). This provides progressive photos of the fluid dynamics of the emitting species as well as a leading edge plume expansion velocity. It is especially useful for hot particulate matter from laser ablation.

Photo-Multiplier Tubes (PMT's) which measure the overall intensity of emitted light can be linearly arrayed with increasing distance from the target surface, while situated to view the ever-expanding plume in succession (TOF Plasma Recombination Light diagnostic) (Bak87). The total light signal from each PMT versus time can then be used to estimate the "center of mass velocity" away from the target and the kinetic temperature of the expanding plume in the center of mass reference frame.

The plume can also be photographed by using an expanded and collimated dye laser beam passing through the ablation plume and then imaged directly onto the film (Shadowgraphy) (Gup91, Sri93). Variations of dark and light on the film denote refraction by a shock wave, an ablation plume density gradient, or nonresonant absorption by dense plume particulate. When the dye laser is tuned to an atomic or molecular resonance, darkness on the film denotes the presence of resonant plume constituents as well. This has been called Absorption Photography (Mos88), Resonant Shadowgraphy (Sie77, Jel81), or Dye-Laser-Resonance-Absorption-Photography (DLRAP) (Gil91, Ven92).

The observation of the light refracted by the ablation plume can be enhanced by blocking the refracted light from reaching the film. For nonresonant laser wavelengths, this is called Schlieren Photography (Ven90, Ven92a), and for resonant laser wavelengths this is called Resonant Schlieren Photography (Jel81, Koo79, Mos88, Sie77).

Finally, interferometric methods may be utilized to observe ablation plumes. This entails splitting the laser beam into two components, expanding and collimating them both, passing one beam through the ablation plume (the object beam), and recombining the two beams to form an interference pattern. The total phase shift of the laser light passing through the ablation plume is measured as a function of location on the recorded interferogram. By taking three interferograms at three different wavelengths, one near a resonant transition and two away from the resonance, the source of the phase shift can be determined and thus give: 1) the electron line-density, 2) the nonresonant atomic or molecular line-density, and 3) the resonant constituent atomic (or ionic) line-density. Of course, the photographic images of the plume expansion provide information concerning the fluid dynamics and the expansion velocities of the three, measurable components.

There are three types of interferometry that can be used to study plume expansion. If the object beam passes through the ablation plume once and the two beams are recombined to be collinear, it is called Mach-Zehnder Interferometry. If the object beam passes through the ablation plume twice by reflection off a mirror and the two beams are recombined to be collinear, it is called Michelson Interferometry. If, 1) the object beam passes through the plume once, 2) the two beams are recombined on a photographic emulsion, and 3) the laser is pulsed once with the ablation plume present, and once without the plume, it is called Holographic Interferometry. As of this writing, one laboratory is known to have utilized nonresonant Michelson interferometry (Wal86) to study laser ablation plume expansion, and two other laboratories are known to have utilized resonant holographic interferometry (Dre75, Koo79).

A derivation and calculation of an extension to the theory of resonant holographic interferometry, for use in the interpretation of resonant holograms at wavelengths very near an atomic resonance, will be presented in Chapter 2. The experimental configuration will be described in Chapter 3. The resonant holograms of the expanding ablation plumes and their interpretation will be presented in Chapter 4, and the overall conclusions of this work will be presented in Chapter 5.

CHAPTER 2 THEORY

Holographic Interferometry

When light is incident upon a neutral gas, it will interact with the gas with some combination of absorption, reflection, refraction, and phase shift. If the absorption is due to a resonant interaction with a quantum mechanical transition in an atomic or molecular component of the gas, then the light will be re-emitted in a random direction with respect to that gas particle. If the gas is relatively transparent, the line-integrated-density of the gas can be determined by measuring the phase shift of a laser beam passing through the gas. With a tunable dye laser, the laser wavelength can be tuned very near a resonant transition of one species of the gas and, if resonant absorption is not excessive, the line-integrated-density of the aforementioned gas species can be measured with several orders of magnitude more sensitivity than with a nonresonant wavelength.

Holographic interferometry provides an accurate, photographic method to measure the incurred phase shift of the laser beam after it has passed through a transparent medium. There are a number of good textbooks that present detailed discussions of holographic interferometry, including Vest (Ves85) and Hecht (Hec79). Figure 2.1 shows a schematic of a holographic interferometry setup. The laser beam is split into two



Figure 2.1 Schematic For Recording The Hologram. BS=Beam Splitter, M=Mirror, and HP=Holographic Plate.



Figure 2.2 Schematic For Reconstructing The Hologram. HP=Holographic Plate, C= Camera.

parts by a beam splitter. Both beams are expanded with a pair of convex lenses and cleaned up with an iris at the focal point. One beam passes through the object of interest before both beams are incident upon the photographic emulsion of a holographic plate. The laser is then pulsed a second time without the test object present. The holographic plate is developed appropriately and then illuminated only by the reference beam, as shown in Figure 2.2. The laser-exposed-and-then-developed portion of the holographic plate forms a diffraction pattern that causes three images to emerge from the holographic plate. The undeflected beam represents the zeroeth order wave and it contains no phase information about the test object. The other two beams are the first order deflections of the holographic diffraction pattern which contain both phase and amplitude information concerning the test object. One of the beams is the image of the test object and the other beam is its conjugate image.

Mathematically, the electric field of the laser beam can be written

$$\underline{E}(\mathbf{x}, \mathbf{y}, \mathbf{z}, \boldsymbol{\omega}, \mathbf{t}_{\mathrm{L}}) = S(\mathbf{x}, \mathbf{y}, \mathbf{t}_{\mathrm{L}}) \exp[i\boldsymbol{\omega}\mathbf{t} - i\boldsymbol{\Phi}(\mathbf{x}, \mathbf{y}, \mathbf{z}, \boldsymbol{\omega})]$$
(2.1)

where (x,y) are the spatial coordinates on the holographic plate, (z) is the direction of propagation of the laser beam, the term $S(x,y,t_L)$ includes the temporal and spatial pulse shape of the laser, and where (dropping the (x,y) notation)

$$\Phi(\omega, z) = \frac{\omega}{c} \int_0^z \tilde{n}(\omega, z') dz' = \frac{\omega z}{c} + \frac{\omega}{c} \int_0^z [\tilde{n}(\omega, z') - 1] dz'$$

$$= \frac{\omega z}{c} + \phi(\omega, z) - iK(\omega, z)$$
, (2.2)

where ϕ and K are real, and thus

$$\underline{\mathbf{E}} = \mathbf{S} \exp\left[i\omega\left(\mathbf{t} - \frac{\mathbf{z}}{\mathbf{c}}\right) - i\phi - \mathbf{K}\right]$$
(2.3)

The intensity of light incident onto the holographic plate during a single laser pulse is

$$I = |\underline{E}_{O} + \underline{E}_{R}|^{2} = S_{O}^{2} \exp(-2K_{O}) + S_{R}^{2} \exp(-2K_{R})$$

$$+ S_{O} S_{R} \exp(-K_{O} - K_{R} - i\phi_{R} + i\phi_{O})$$

$$+ S_{O} S_{R} \exp(-K_{O} - K_{R} + i\phi_{R} - i\phi_{O})$$

$$= S_{O}^{2} \exp(-2K_{O}) + S_{R}^{2} \exp(-2K_{R})$$

$$+ 2S_{O} S_{R} \exp(-K_{O} - K_{R}) \cos(\phi_{R} - \phi_{O})$$

$$(2.4)$$

where the subscripts O and R stand for the object and reference beams, respectively. When this single-pulsed hologram is reconstructed, the transmission of light immediately on the right side of the holographic plate in Figure 2.2 is

$$\underline{\mathbf{t}} = \mu_{o} \underline{\mathbf{E}}_{C} \left(\left| \underline{\mathbf{E}}_{O} \right|^{2} + \left| \underline{\mathbf{E}}_{R} \right|^{2} \right) + \theta_{o} \underline{\mathbf{E}}_{C} \underline{\mathbf{E}}_{O}^{*} \underline{\mathbf{E}}_{R} + \theta_{o} \underline{\mathbf{E}}_{C} \underline{\mathbf{E}}_{O} \underline{\mathbf{E}}_{R}^{*} , \qquad (2.6)$$

where μ_0 is an amplitude transmission constant with respect to the photographic developing process for the zeroeth order term and θ_0 is a similar photographic constant for the first order terms. A photograph of the third term in Equation 2.6 does not record any information about the phase of the test object, since the laser was only pulsed once. When the dye laser is pulsed twice, once with the test object present and once without, the two reconstructions of the test object propagate collinearly and the irradiance on the film of the camera in Fig. 2.2 is

$$I_{c} = \left| \theta_{o} \underline{E}_{C} (\underline{E}_{O1} \underline{E}_{R1}^{*} + \underline{E}_{O2} \underline{E}_{R2}^{*}) \right|^{2} , \qquad (2.7)$$

which gives

$$I_{C} = \theta_{o}^{2} |\underline{E}_{C}|^{2} S_{O}^{2} S_{R}^{2} \left\{ \exp \left[-2 K_{R1} - 2 K_{O1} \right] + \exp \left[-2 K_{R2} - 2 K_{O2} \right] + 2 \exp \left[-K_{R1} - K_{O1} - K_{R2} - K_{O2} \right] \cos \left(\phi_{O1} - \phi_{O2} - \phi_{R1} + \phi_{R2} \right) \right\}$$
(2.8)

If the phase front of the reference beam is fairly constant over the width of the expanded laser beam and if absorption is minimal, the $\cos(\phi_{01}-\phi_{02}-\phi_{R1}+\phi_{R2})$ term dominates the black and white pattern on the film. Also, any imperfections in the system optics are canceled out in the cosine term. If the mirror in the reference beam is turned slightly in between laser pulses, straight fringes appear on the reconstruction and any deviation from straight is then due to the test object. The amount of deviation is used to determine the absolute-line-integrated-density of the gas species being investigated.

Resonant Holographic Interferometry

For wavelengths very close to a resonant atomic or molecular transition of a component gas species, the real part of the index of refraction is written as (Gri64, Dre75, Ost75, Ost85)

$$\operatorname{Re}[\tilde{n}(\omega, z) - 1] = \frac{e^{2} N_{i}(z) f_{o}}{4m\omega_{o}\varepsilon_{o}} \left[\frac{\omega_{o} - \omega}{(\omega - \omega_{o})^{2} + \frac{\gamma^{2}}{4}} \right]$$
(2.9)

where $N_i(z)$ is the resonant particle density, f_0 is the oscillator strength of the transition with resonant frequency ω_0 , and γ is the dissipative damping constant. This allows the total resonant phase shift to be calculated using the Equation

$$\left(\frac{\Delta s}{s}\right) = \frac{\Delta \phi_{\text{resonant}}}{2\pi} = \frac{\omega_o}{2\pi c} \int_{o}^{z} \text{Re}[\tilde{n}(\omega, z') - 1] dz' \qquad (2.10)$$
Equation 2.9 assumes $|\omega - \omega_0|$ is at least several times larger than $\gamma/2$ and that Doppler broadening effects can be ignored. For $|\omega - \omega_0| \approx \gamma/2$, Equation 2.9 can be written as (Dre71, Dre73, Bac80, Bow85, Rub91)

$$\operatorname{Re}[\tilde{n}(\omega, z) - 1] = \frac{e^2 N_i(z) f_o}{4m\omega_o \varepsilon_o} \left[\frac{1}{\omega_o - \omega} \right] , \qquad (2.11)$$

which is called Sellmeier's Equation.

Equation 2.9 can be derived in one of two ways. The first method begins with the Kramers-Kronig Equations (Gri64, Koo78, Cra86, Mos88)

$$\operatorname{Re}[\tilde{n}-1] = \frac{1}{\pi} \operatorname{PV} \int_{-\infty}^{\infty} \frac{\operatorname{Im}[\tilde{n}-1]}{\omega' - \omega} d\omega'$$
(2.12)

and

$$Im[\tilde{n}-1] = \frac{1}{\pi} PV \int_{-\infty}^{\infty} \frac{Re[\tilde{n}-1]}{\omega'-\omega} d\omega'$$
(2.13)

where PV stands for Cauchy's principle value, and then assumes that

$$\operatorname{Im}[\tilde{n}-1] = \frac{e^2 N f_o}{4 m \omega_o \varepsilon_o} \left[\frac{\gamma/2}{(\omega - \omega_o)^2 + \frac{\gamma^2}{4}} \right]$$
(2.14)

Evaluating Equation 2.12 with Equation 2.14 then gives Equation 2.9. Together, ñ-1 can then be written (Gri64)

$$\tilde{n} - 1 = \frac{e^2 N_i f_o}{4m\omega_o \varepsilon_o} \left[\frac{-1}{\omega - \omega_o + i\frac{\gamma}{2}} \right]$$
(2.15)

The second method to derive Equation 2.9 is to assume that the transparent test section can be treated as a classical dielectric medium, thus giving (Mar67, Mea70, Ber71, Sie77, Koo79, Ode80, Jel81)

$$\frac{\tilde{n}^2 - 1}{\tilde{n}^2 + 2} = \frac{1}{3\varepsilon_0} \sum_i \tilde{\alpha}_i N_i$$
(2.16)

and then use the quantum mechanical value of the complex polarizability

$$\tilde{\alpha}_{i} = \frac{\begin{pmatrix} e^{2}/m \end{pmatrix} f_{ik}}{\omega_{ik}^{2} - \omega^{2} + i\gamma_{ki}\omega}$$
(2.17)

Equation 2.9 (as well as Equation 2.15) follows when assuming that only one resonance transition is significant. For a more detailed derivation of the second method, see Appendix 1.

When $|\omega - \omega_0| \approx \gamma/2$ and when Doppler broadening is significant in comparison to natural line broadening (and other broadening mechanisms), Equation 2.9 is modified by using the Doppler shifted frequency (Mar67, Mea70, Jel81, Küg83, Cra86, Bon86, Woi91, Woi92)

$$\omega' \equiv \omega \left(1 + \frac{v_x}{c} \right) , \qquad (2.18)$$

where v_x is the velocity component in the direction of the laser beam, and by using a Maxwell-Boltzmann distribution of velocities for the gaseous test section

$$N'_{i} = N_{i} \left(\frac{m}{2\pi k_{B}T}\right)^{1/2} \int_{-\infty}^{\infty} dv_{x} \exp\left(\frac{-mv_{x}^{2}}{2k_{B}T}\right)$$
(2.19)

Using Measures' notation, this leads to (see Appendix 1)

$$\tilde{n} - 1 \cong \left\{ \frac{c}{\omega_{o}} \right\} (-iA) N_{i} W(-u + i\alpha) , \qquad (2.20)$$

or when line-integrating along the path of the laser beam

$$\phi - ik = -i A \overline{N_i L} W(-u + i\alpha)$$
(2.21)

where

$$\beta = \frac{\omega_o}{c} \sqrt{\frac{2k_BT}{m}} , \quad \alpha = \frac{\gamma_{ik}}{2\beta} , \quad u = \frac{\omega - \omega_o}{\beta} , \quad A = \left(\frac{\pi^{3/2} r_o c f_{ik}}{\beta}\right) , \quad (2.22)$$

 $\overline{N_iL}$ is the line-integrated-density, $r_o = e^2 / (4\pi\epsilon_o mc^2) = 2.818 \times 10^{-13}$ cm is the classical electron radius and

$$W(z) = \exp(-z^2) \times (1 + erf\{iz\})$$
(2.23)

where the real and imaginary parts of W(z) for Im[z]=0 are shown in Figure 2.3. Noting that for $|z|\gg 2$ (Abr72)

$$W(z) \rightarrow \frac{iz}{\sqrt{\pi} |z|^2} , \qquad (2.24)$$

equation 2.15 is recovered.

It is notable that the method for accounting for Doppler broadening used above is also called the Voigt profile method since it begins with a



Fig. 2.3 The basic response curve for W(z), and thus for resonant interferometry when accounting for Doppler Shift.

Lorentzian-shaped natural broadening, Equation 2.15, and convolutes it with a Gaussian-shaped velocity distribution/line broadening mechanism, Equation 2.19.

Another notable item is the polarity of Im[fi-1] since there is some discrepancy in the open literature on this point. In Chapter 4, experimental evidence is presented to agree with (Gri64, Mar67, Mea70, Dre71, Ber71, Dre73, Dre75, Ost75, Sie77, Koo79, Koo78, Jel81, Ost85, Bow85, Cra86, Mos88, Sta88) that

$$\operatorname{sgn}\left\{\operatorname{Im}\left[\tilde{n}-1\right]\right\} = -\operatorname{sgn}\left\{\omega-\omega_{o}\right\}$$
(2.25)

is correct (as in Fig. 2.3), where $sgn{x}$ is the function

$$sgn(x) = \begin{cases} 1 & x > 0 \\ 0 & x = 0 \\ -1 & x < 0 \end{cases}$$
, (2.26)

as opposed to the polarity

$$\operatorname{sgn}\left\{\operatorname{Im}\left[\tilde{n}-1\right]\right\} = +\operatorname{sgn}\left\{\omega-\omega_{o}\right\}$$
(2.27)

given by (Küg83, Sei86, Woi91, Woi92). In addition, Odenthal, et al. (Ode80) and Rubin, et al. (Rub91) used Equations following Eqn. 2.25 but presented graphs following Eqn. 2.27.

Accounting For Finite Laser Bandwidth

Dreiden, et al. (Dre75) accounted for finite laser bandwidth by ignoring Doppler broadening (and thus using Eqn. 2.9), assuming that the line-integrated-density of the test section was near the lower detectability limit, and by doing a simple average of the calculated fringe shift over the FWHM of the laser line,

$$\overline{\left(\frac{\Delta s}{s}\right)} = \frac{\overline{\mathrm{NLr}}_{o} cf_{o}}{2 \,\delta \omega} \int_{\omega - \left(\frac{\delta \omega/2}{2}\right)}^{\omega + \left(\frac{\delta \omega/2}{2}\right)} d\omega_{1} \frac{\omega_{0} - \omega_{1}}{\left(\omega_{1} - \omega_{0}\right)^{2} + \left(\frac{\gamma}{2}\right)^{2}}$$

$$= \frac{\overline{\mathrm{NLr}}_{o} cf_{o}}{2 \,\delta \omega} \frac{1}{2} \ln \left(\frac{1 + x}{1 - x}\right) = \frac{\overline{\mathrm{NLr}}_{o} cf_{o}}{2 \,\delta \omega} \arctan(x)$$
(2.28)

where

$$x = \frac{\left(\omega_0 - \omega_1\right)\delta\omega}{\left(\omega_1 - \omega_0\right)^2 + \left(\gamma/2\right)^2 + \left(\delta\omega/2\right)^2}$$
(2.29)

and $\delta \omega$ is the FWHM bandwidth of the laser intensity. With this analysis, Dreiden determined that an increase in laser bandwidth causes a decrease in the sensitivity of the diagnostic and a contraction of the range of measurable line-densities using resonant interferometry techniques.

To more fully account for the interaction between the Gaussian shape of the laser bandwidth used in this work and the Voigt profile obtained for the region $|\omega - \omega_0| \ge \approx (\gamma/2)$ when accounting for Doppler broadening, and to account for the interference effects induced at the holographic plate by a finite bandwidth light source, the temporalcoherence discussions of Goodman for interferometry are utilized (Goo85). Here, the laser is simplistically modeled as producing one photon at a time with a random frequency, ω , which has a Gaussian probability distribution about a center frequency, ω . It is then sufficient to time-average the random frequency of the laser. By convention, this is written

$$\underline{\Gamma}(\tau) = \left\langle \underline{E}(t+\tau)\underline{E}^{*}(t) \right\rangle = \lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{+T/2} \underline{E}(t+\tau)\underline{E}^{*}(t)dt \qquad (2.30)$$

where $\underline{\Gamma}(\tau)$ is the complex autocorrelation function, $\underline{E}(t)$ is the complex electric field of the laser, and < > denotes a time average. The electric

field of the laser is considered ergodic when the time average of $\underline{E}(t + \tau)\underline{E}^{*}(t)$ is equal to the probability average of $\underline{E}(t + \tau)\underline{E}^{*}(t)$; i.e., when

$$\left\langle \underline{E}(t+\tau)\underline{E}^{*}(t)\right\rangle = \overline{\underline{E}(t+\tau)\underline{E}^{*}(t)}$$
(2.31)

where

$$\overline{\underline{E}(t+\tau)\underline{E}^{*}(t)} = \iint_{-\infty}^{\infty} \underline{E}(t+\tau)\underline{E}^{*}(t) p(\underline{E}(t+\tau), \underline{E}^{*}(t); t, t+\tau) d\underline{E} d\underline{E}^{*}$$
(2.32)

and $p(\underline{E}(t+\tau), \underline{E}^{*}(t); t, t+\tau)$ is the probability distribution of the electric fields. If the electric field of the laser is ergodic, then the Wiener-Khinchin Theorem can be used, which states

$$\underline{\Gamma}(\tau) = \int_{-\infty}^{\infty} G(\omega) \exp(-i\omega\tau) d\omega$$
(2.33)

where $G(\omega)$ is the laser frequency probability

$$G(\omega) = \lim_{T \to \infty} \frac{\underline{\underline{\varepsilon}(\omega)} \underline{\underline{\varepsilon}^*(\omega)}}{T}$$
(2.34)

and $\underline{e}(\omega)$ is the Fourier transform of the complex electric field

$$\underline{\varepsilon}(\omega) = \int_{-\infty}^{\infty} \underline{E}(t) \exp(-i\omega t) dt \qquad (2.35)$$

Given

$$\underline{\Gamma}(\tau=0) = \int_{-\infty}^{\infty} G(\omega) \, d\omega = S^2$$
(2.36)

and

$$\hat{G}(\omega) = \frac{G(\omega)}{S^2} , \qquad (2.37)$$

the laser frequency probability with a Gaussian frequency distribution is then written as

$$\hat{G}(\omega) \cong \frac{2\sqrt{\ln 2}}{\sigma\sqrt{\pi}} \exp\left[-\left(2\sqrt{\ln 2}\frac{\omega-\varpi}{\sigma}\right)^2\right], \qquad (2.38)$$

where σ is the Full Width at Half Maximum (FWHM) of the laser intensity. This means that the time average of the electric fields on the holographic plate can be written as

$$\underline{\Gamma}(\tau) = \frac{2S^2 \sqrt{\ln 2}}{\sigma \sqrt{\pi}} \int_{-\infty}^{\infty} \exp\left[-4\ln 2\left(\frac{\omega - \overline{\omega}}{\sigma}\right)^2 - i\omega\tau\right] d\omega \qquad (2.39)$$

Note that Eqn. 2.39 only includes temporal coherence (two beams delayed in time) and not spatial coherence (two beams shifted in space). For this work, it is assumed that the spatial coherence of the laser is sufficient so that the relevant mathematics may be neglected.

Equation 2.4 can now be rewritten

$$\langle I \rangle = \Gamma_{OO} + \Gamma_{RR} + \Gamma_{OR} + \Gamma_{OR}^*$$
(2.40)

and Eqn. 2.7 can be rewritten as

$$\langle I_{c} \rangle = \left| \theta_{o} \underline{E}_{c} (\Gamma_{OR1} + \Gamma_{OR2}) \right|^{2}$$
 (2.41)

where "1" and "2" refer to the first and second dye laser pulses, respectively.

The Solution for Doppler Broadening and Finite Bandwidth

Recalling the notation for the electric fields given in Equations 2.1-2.3, the two laser beams for both laser pulses can be written

$$\Phi_{01} = \frac{\omega z}{c} + \omega \tau_0 - i K_0 \qquad (2.42)$$

$$\Phi_{R1} = \frac{\omega z}{c} + \omega \tau_R - i K_R \qquad (2.43)$$

$$\Phi_{O2} = \frac{\omega z}{c} + \omega \tau_O - iK_O - iA \overline{N_i L} W(-u + i\alpha) , \qquad (2.44)$$

and

$$\Phi_{R2} = \frac{\omega z}{c} + \omega \tau_R - iK_R + \omega y \delta , \qquad (2.45)$$

where K_O and K_R represent any absorption incurred outside the test section by the object and reference beams, respectively, and t_O and t_R refer to the travel-time of the object and reference beams, respectively, as they travel between the incident side of the f and the holographic plate (see Fig. 2.1).

The Doppler broadened test section $(-iA \overline{N_iL} W(-u+i\alpha))$ is now included in Φ_{O2} and the reference fringes caused by turning a mirror in the reference beam between laser pulses, $\omega y \delta$, is accounted for in Φ_{R2} . This notation for reference fringes means that the black and white stripes are in the \hat{x} direction with the variation in the \hat{y} direction whose spacing is determined by the constant δ . Recall also that each of these variables is a function of (x,y), the spatial coordinates on the holographic plate; their notation is omitted here for simplicity.

In anticipation of using Eqns. 2.42-2.45 in Eqn. 2.39, one writes

$$i\omega \tau_{001} = i(\Phi_{01} - \Phi_{01}^*) = 2K_0$$
, (2.46)

$$i\omega \tau_{RR1} = i(\Phi_{R1} - \Phi_{R1}^{*}) = 2K_{R}$$
, (2.47)

$$i\omega\tau_{OR1} = i(\Phi_{O1} - \Phi_{R1}^{*}) = i\omega(\tau_{O} - \tau_{R}) + K_{O} + K_{R}, (2.48)$$

$$i\omega \tau_{002} = i \left(\Phi_{02} - \Phi_{02}^* \right)$$
 (2.49)

$$= 2K_0 + A \overline{N_iL} \left[W(-u + i\alpha) + W(+u + i\alpha) \right]$$

(utilizing the relation (Abr72)

$$[W(z)]^* = W(-[z]^*) \}, \qquad (2.50)$$

$$i\omega \tau_{RR2} = i \left(\Phi_{R2} - \Phi_{R2}^* \right) = 2 K_R$$
, (2.51)

and

$$i\omega\tau_{OR2} = i\left(\Phi_{O2} - \Phi_{R2}^{*}\right)$$
(2.52)

$$= i\omega \left(\tau_{O}^{} - \tau_{R}^{} - \delta y\right) + K_{O}^{} + K_{R}^{} + A \overline{N_{i}L} W(-u + i\alpha).$$

Now inserting Equations 2.46, 2.47, 2.48, and 2.51 into Equation 2.39, one obtains

$$\underline{\Gamma}_{OO1} = \underline{\Gamma}(\tau_{OO1}) = S_O^2 \exp[-2K_O] , \qquad (2.53)$$

$$\underline{\Gamma}_{RR1} = S_R^2 \exp[-2K_R] , \qquad (2.54)$$

$$\underline{\Gamma}_{OR1} = S_O S_R \exp\left[-K_O - K_R - \frac{\sigma^2 (\tau_O - \tau_R)^2}{16 \ln 2} - i \overline{\omega} (\tau_O - \tau_R)\right], \quad (2.55)$$

and

$$\underline{\Gamma}_{RR2} = S_R^2 \exp[-2K_R]$$
(2.56)

The $\sigma^2(\tau_0 - \tau_R)^2/(16\ln 2)$ term in Equation 2.55 is the temporal coherence term. If the time-travel difference between the object and reference beams becomes much larger than the coherence-time of the laser, $(2\pi/\sigma)$, then $\underline{\Gamma}_{OR1} \rightarrow 0$ and no fringes are observed at all.

 $\underline{\Gamma}_{OO2}$ and $\underline{\Gamma}_{OR2}$ are not analytically solvable using Equation 2.39 because of the presence of W(z) in the exponential, so their solutions are approximated in this work using the Saddle Point Theorem (Arf85), where

$$\int_{-\infty}^{\infty} dx \exp[f(x)] \cong \int_{-\infty}^{\infty} dx \exp\left[f(x_s) + \frac{1}{2}f''(x_s)\{x - x_s\}^2\right]$$
(2.57)

$$\cong \eta \exp[f(\mathbf{x}_{\mathrm{g}})] \sqrt{\frac{1}{f''(\mathbf{x}_{\mathrm{g}})}}$$

where η is a constant and x_s is a solution of $f'(x_s) \equiv 0$.

Inserting Equation 2.49 into Equation 2.39 gives

$$\underline{\Gamma}_{OO2} = \frac{2S_O^2 \sqrt{\ln 2}}{\sigma \sqrt{\pi}} \int_{-\infty}^{\infty} d\omega \exp\left[-4\ln 2\left(\frac{\omega - \varpi}{\sigma}\right)^2 - 2K_O\right]$$

$$-A \overline{N_i L} \left[W(-u + i\alpha) + W(+u + i\alpha)\right]$$
(2.58)

Letting $\omega' = \omega - \omega$, $\Delta = (\omega - \omega_0)/\beta$, $\Lambda = \omega'/\beta$, and $u = \Delta + \Lambda$, Equation 2.58 becomes

$$\underline{\Gamma}_{002} = \frac{2S_0^2 \beta \sqrt{\ln 2}}{\sigma \sqrt{\pi}} \int_{-\infty}^{\infty} d\Lambda \exp[h(\Lambda)]$$
(2.59)

where

$$h(\Lambda) = -4\ln 2\left(\frac{\beta}{\sigma}\right)^2 \Lambda^2 - 2K_0 - A \overline{N_i L} \left[W(\psi) + W(\chi)\right] , \qquad (2.60)$$

 $\psi = -\Delta - \Lambda + i\alpha$, and $\chi = +\Delta + \Lambda + i\alpha$. (Note that $\text{Im}[W(\psi) + W(\chi)] = 0$.) This in turn leads to

$$h'(\Lambda) = -8\ln 2\left(\frac{\beta}{\sigma}\right)^2 \Lambda - 2A \overline{N_i L} \left[\psi W(\psi) + \chi W(\chi)\right]$$
(2.61)

and

$$h''(\Lambda) = -8\ln 2 \left(\frac{\beta}{\sigma}\right)^2 - 2A \overline{N_i L} \left\{ W(\psi) [2\psi^2 - 1] - W(\chi) [2\chi^2 - 1] - \frac{4i}{\sqrt{\pi}} (\Delta + \Lambda) \right\}, \qquad (2.62)$$

where Equation 2.61 is solved for the complex solution $\underline{\Lambda_s}$ from $h'(\underline{\Lambda_s}) \equiv 0$ using Newton's Method in Mathematica (see Appendix 2). Note that as $\sigma \to 0$ in Eqn. 2.62 and thus $\underline{\Lambda_s} \to 0$ in the solution to $h'(\underline{\Lambda_s}) = 0$, Eqn. 2.58 becomes solvable analytically, giving

$$\underline{\Gamma}_{002} \rightarrow S_0^2 \exp\left[-2K_0 - A \overline{N_i L} \left\{W(\psi) + W(\chi)\right\}\right]$$
(2.63)

Therefore for small Λ_s ,

$$\eta_{\rm h} \approx \sqrt{\frac{\pi}{4\ln 2}} \left(\frac{\sigma}{\beta}\right) \sqrt{h''(0)} \tag{2.64}$$

0ľ

$$\underline{\Gamma}_{002} \cong S_0^2 \sqrt{\frac{h''(0)}{h''(\Lambda_s)}} \exp\left[h(\Lambda_s)\right]$$
(2.65)

Although $h''(\Lambda)$ is complex-valued, it was numerically determined here that $(h''(0)/h''(\Lambda)) \approx |h''(0)/h''(\Lambda)|$ and thus justifying the use of the absolute value in Eqn. 2.65.

Finally, Equation 2.52 and Equation 2.39 combine to make

$$\underline{\Gamma}_{OR2} = \frac{2S_O S_R \beta \sqrt{\ln 2}}{\sigma \sqrt{\pi}} \int_{-\infty}^{\infty} d\omega \exp \left[-4 \ln 2 \left(\frac{\omega - \overline{\omega}}{\sigma} \right)^2 - K_O - K_R -i\omega (\tau_O - \tau_R - \delta y) - A \overline{N_i L} W(-u + i\alpha) \right] .$$
(2.66)

Letting $\omega' = \omega - \overline{\omega}$, $\Delta = (\overline{\omega} - \omega_o)/\beta$, $\varepsilon = \omega'/\beta$, $u = \Delta + \varepsilon$, and $\Theta = -\Delta - \varepsilon + i\alpha$, Equation 2.66 becomes

$$\underline{\Gamma}_{OR2} = \frac{2S_O S_R \beta \sqrt{\ln 2}}{\sigma \sqrt{\pi}} \int_{-\infty}^{\infty} d\varepsilon \exp[g(\varepsilon)]$$
(2.67)

where

$$g(\varepsilon) = -4\ln 2\left(\frac{\beta}{\sigma}\right)^2 \varepsilon^2 - K_0 - K_R - i\varpi(\tau_0 - \tau_R - \delta y) - A \overline{N_i L} W(\Theta)$$
(2.68)

$$g'(\varepsilon) = -8\ln 2\left(\frac{\beta}{\sigma}\right)^2 \varepsilon - 2A \ \overline{N_i L} \ \Theta \ W(\Theta) + i\left(\frac{2}{\sqrt{\pi}}\right)A \ \overline{N_i L}$$
(2.69)

and

$$g''(\varepsilon) = -8\ln 2\left(\frac{\beta}{\sigma}\right)^2 - 2A \overline{N_iL} \quad W(\Theta) \left\{2\Theta^2 - 1\right\} + i\left(\frac{4}{\sqrt{\pi}}\right)A \overline{N_iL} \Theta \quad (2.70)$$

The $\omega'(\tau_0 - \tau_R - \delta y)$ term has been dropped to make the integration possible, thereby assuming $\omega' \ll \overline{\omega}$ for the values of ω' where the 4 ln2 $(\omega/\sigma)^2$ term does not dominate in the exponential of Eqn. 2.66. Equation 2.69 is solved for the complex solution $\underline{\varepsilon}_g$ from $g'(\underline{\varepsilon}_g) \equiv 0$ by rearranging it to the form

$$W(-\Delta - \underline{\varepsilon}_{s} + i\alpha) = \frac{-\left(\frac{4\ln 2\beta^{2}}{A \overline{N_{i}L}\sigma^{2}}\right)\underline{\varepsilon}_{s} + i\left(\frac{1}{\sqrt{\pi}}\right)}{\left(-\Delta - \underline{\varepsilon}_{s} + i\alpha\right)}$$
(2.71)

and then solving for $\underline{\epsilon}_{s}$ using Newton's Method in Mathematica (see Appendix 2). Again noting that as $\sigma \to 0$ in Eqn. 2.66 and thus $\underline{\epsilon}_{s} \to 0$ in the solution to $g'(\underline{\epsilon}_{s}) \equiv 0$, Eqn. 2.66 becomes solvable analytically, giving

$$\frac{\Gamma_{OR2} \rightarrow S_O S_R \exp\left[-K_O - K_R - \frac{\sigma^2 (\tau_O - \tau_R - \delta y)^2}{16 \ln 2} -i\varpi (\tau_O - \tau_R - \delta y) - A \overline{N_i L} W(-\Delta + i\alpha)\right]$$
(2.72)

Therefore for small $\frac{\varepsilon_s}{s}$,

$$\eta_g \approx \sqrt{\frac{\pi}{4\ln 2}} \left(\frac{\sigma}{\beta}\right) \sqrt{g''(0)} \exp\left[-\frac{\sigma^2 (\tau_O - \tau_R - \delta y)^2}{16\ln 2}\right]$$
(2.73)

or

$$\underline{\Gamma}_{OR2} \cong S_O S_R \sqrt{\frac{g''(0)}{g''(\frac{\varepsilon}{s})}} \exp\left[-\frac{\sigma^2(\tau_O - \tau_R - \delta y)^2}{16 \ln 2} + g(\frac{\varepsilon}{s})\right] (2.74)$$

Two roots were found for Eqn. 2.71. One of the roots changed continuously as Δ and NL were varied, while the other root was discontinuous for the same parameter regions. Therefore the first root was used in the solutions that follow.

Equation 2.41 gives the irradiance onto the reconstruction camera film. Assuming the temporal coherence requirement is easily met,

$$\frac{\sigma^2 (\tau_O - \tau_R - \delta y)^2}{16 \ln 2} \approx \frac{\sigma^2 (\tau_O - \tau_R)^2}{16 \ln 2} \approx \text{very small} , \qquad (2.75)$$

Equation 2.41 can now be written

$$\langle I_c \rangle = M \exp[-N] \{ 1 + P + Q \cos[T] \}$$
 (2.76)

where

$$M = \theta_o^2 |\underline{E}_C|^2 S_O^2 S_R^2$$
(2.77)

$$N = 2K_{O} + 2K_{R} + \frac{\sigma^{2}(\tau_{O} - \tau_{R})^{2}}{8\ln 2}$$
(2.78)

$$P = \Omega_g^2 \exp\left[2R\left(\underline{\varepsilon_s}\right)\right]$$
(2.79)

$$Q = 2\Omega_{g} \exp\left[R\left(\underline{\varepsilon}_{s}\right)\right] = 2\sqrt{P}$$
(2.80)

$$T = \varpi y \delta + Y\left(\frac{\varepsilon}{\underline{s}}\right)$$
 (2.81)

and where

$$\Omega_{g} = \sqrt{\left|\frac{g''(0)}{g''(\underline{e}_{s})}\right|}$$
(2.82)

$$R(\underline{\varepsilon_{s}}) = Re\left[-4\ln 2\left(\frac{\beta}{\sigma}\right)^{2}\underline{\varepsilon_{s}}^{2} - A \overline{N_{i}L} W(-\Delta - \underline{\varepsilon_{s}} + i\alpha)\right]$$
(2.83)

$$Y(\underline{\varepsilon_{s}}) = Im\left[-4\ln 2\left(\frac{\beta}{\sigma}\right)^{2}\underline{\varepsilon_{s}}^{2} - A \overline{N_{i}L} W(-\Delta - \underline{\varepsilon_{s}} + i\alpha)\right] .$$
(2.84)

Since the reference fringes, $\varpi_y \delta$, are manually removed when the reconstructed holograms are analyzed, that term may be dropped here and thus the fringe shift can be interpreted as

$$\left(\frac{\Delta s}{s}\right) = \frac{T}{2\pi} = \frac{Y(\varepsilon_s)}{2\pi}$$
(2.85)

Notice that as $\sigma \to 0$ in Eqn. 2.66 and thus $\underline{e_s} \to 0$ in the solution to $g'(\underline{e_s}) \equiv 0$, Measures' answer is obtained (Mea70)

$$\left(\frac{\Delta s}{s}\right) \rightarrow \operatorname{Im}\left[-\frac{A \ \overline{N_{i}L}}{2\pi} W(-\Delta + i\alpha)\right]$$
 (2.86)

The ablation plumes in this work are assumed to have a Maxwellian parallel velocity distribution (i.e., as seen by the dye laser beam) in the range 0.1 eV - 1.0 eV (Dre86, Dre91). As an example of fringe shift interpretation for holographic interferometry, Measures' theory (Eqn. 2.86) is graphed in Fig. 2.4 using the temperatures 0.1 eV, 0.3 eV, and 1.0 eV for the Al(I), 394.401 nm, resonance line ($f_0 = 0.115$). Although the diagnostic is most sensitive for $|\Delta\lambda| \approx 0.0025$ nm, that is also the region of most uncertainty due to the plume temperature. However for $|\Delta\lambda| > 0.010$ nm, the fringe shift interpretation is no longer dependent upon temperature. Since the temperatures of the ablation plumes in this work were not measured, a temperature of 0.3 eV was assumed for the interpretation of all holograms.

Figure 2.5 compares the Gaussian dye laser bandwidth to Measures' interpretation curves for four temperatures. Note that as the temperature increases, the maximum amount of observable fringe shift decreases, the wavelength of the peak fringe shift increases, and the effective bandwidth of the dye laser narrows.

Whereas line-density versus the fringe shift is a linear relationship at each $\Delta\lambda$ in Fig. 2.4 and Fig. 2.5, this is not the case for the theory developed here. Table 2.1 shows example values of the parameters used in this work, and Figure 2.6 shows a sample hologram interpretation graph at



Fig. 2.4 Fringe shift per line density, $(\Delta s/s)/NL$, as a function of $\Delta \lambda = (\lambda - \lambda_0)$ for the neutral aluminum line $\lambda_0 = 394.401$ nm. Calculations made used Measures' theory and thus takes Doppler shift into account but assumes an infinitesimal laser bandwidth (Mea70).

 $|\Delta\lambda| = 0.003$ nm as calculated by Griem (Eqn. 2.9), Measures (Eqn. 2.86), Dreiden (Eqn. 2.28), and this work (Eqn. 2.85) ($\delta\lambda = 0.0031$ nm for dye laser). The fringe shift interpretation curves for a selection of other $|\Delta\lambda|$'s is given in Appendix 3. Notice that the value of Eqn. 2.85 is the same as Measures' value for lower fringe shifts, while it increases to 77% larger than Measures' value at 8 fringe shifts. For $|\Delta\lambda| = 0.010$ nm though, the difference in the four interpretation values at 8 fringe shifts is less than 3%, making any of the equations valid for 0.3 eV.



Fig. 2.5 Comparison of the Dye Laser Bandwidth and the Resonant Hologram Interpretation Curves

For the dye laser used in this work, the intensity of the laser beam has a Gaussian-shaped wavelength spread with a FWHM of 0.0031 nm, however there are theoretically at least a few photons of all wavelengths. The higher the line-density, the more absorption of photons with wavelengths near the center of the resonance, and thus the average $|\Delta\lambda|$ that passes through the plume increases with line-density. In turn, a larger $|\Delta\lambda|$ means a smaller fringe shift as observed on the hologram.

Countering this effect is the Doppler shift caused by the energetic plume, as seen by the dye laser beam. Doppler shift causes the laser bandwidth to effectively broaden in the reference frame of the ablated particles. This more heavily weights the small number of photons with

Table 2.1 Example Parameter Values

```
Input:

Al(I), 394.401 nm

f = 0.115

\gamma = 1.47 \times 10^8 Hz

\Delta \lambda = 0.003 nm

T = 0.3 eV

σ = 1.60x10<sup>10</sup> rad/sec

(\delta \lambda = 0.0031 nm)

NL = 4x10<sup>14</sup> cm<sup>-2</sup>
```

<u>Output:</u> $\Delta = 1.554$ $\alpha = 0.00314$ $\beta = 2.34 \times 10^{10}$ rad/sec $A = 2.314 \times 10^{-13}$ cm⁻² $\epsilon = 0.711 - 0.658*I$ F.S. = 4.43



Fig. 2.6 Fringe shift interpretation at 0.003 nm for the Al(I), 394.401 nm line assuming a plume temperature of 0.3 eV.

wavelengths near the peak of the curve in Fig. 2.5, and thus the observed fringe shift increases (when ignoring absorption).

For example, notice in Fig. 2.6 ($|\Delta\lambda| = 0.003$ nm) that for a line density of $4x10^{14}$ cm⁻², Griem predicts an observed fringe shift (FS) of \approx 5.3. Measures predicts a higher value, \approx 7.0 FS, when accounting for the laser bandwidth broadening caused by Doppler shift while assuming an infinitesimal laser bandwidth to begin with and ignoring absorption. On the other hand, Dreiden predicts a lower value, ≈ 4.3 FS, when accounting for the experimental laser bandwidth but ignoring absorption and Doppler shift; i.e., he assumed the plume was cold. Fig. 2.5 implies that a cold plume (T < 0.1 eV) would give less fringe shift at $|\Delta\lambda| = 0.003$ nm than a 0.3 eV plume because the peak sensitivity of the diagnostic would be at $|\Delta\lambda| < 0.001$ nm. Finally, this work predicts a value of ~4.4 FS, which takes into account the experimental laser bandwidth, the Doppler shift effects of a hot plume, and the selective absorption of the laser beam as a function of wavelength. In light of this, Fig. 2.6 shows that when considering the ablation plume to be hot, Doppler shift effects dominate at lower line densities and selective absorption effects dominate at higher line-densities.

To understand the temperature dependance of the hologram interpretation calculations, Measures' interpretation is graphed in Fig. 2.7(a) for four temperatures and $|\Delta\lambda| = 0.003$ nm, and this work's interpretation is similarly graphed in Fig. 2.7(b). (Similar graphs for other $|\Delta\lambda|$'s are shown in Appendix 3.) The temperature dependence is more pronounced in Measures' calculation because the laser bandwidth is infinitesimal, while the temperature dependence is averaged out between 0.1 eV and 1.0 eV due to the laser bandwidth in this work's calculation.







Fig.2.7(b) This work's calculation at 0.003 nm for several ablation plume temperatures.

Therefore, the fringe shift interpretation for the plumes in this work is not very sensitive to temperature since they are assumed to have parallel expansion velocities in the temperature range 0.1 eV - 1.0 eV (Dre86, Dre91).

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CHAPTER 3 EXPERIMENTAL CONFIGURATION

The Main Optical Table

Figure 3.1(a) shows the basic experimental configuration for the optics table utilized for Resonant Holographic Interferometry (RHI) and Dye Laser Resonant Absorption Photography (DLRAP). A Lumonics Excimer Laser (model TE-292-K,< 0.8 J, 40 ns, 0.5 Hz max.) filled with a premixed KrF laser mixture (248 nm output wavelength) is focused onto a solid aluminum target inside a glass, six-way, vacuum chamber, thus producing an ablation plume. To probe the ablation plume, a Lambda Physik XeCl excimer-pumped dye laser (excimer model LPX-105, dye laser model FL-3002E, 20 ns, \approx 5 mJ, 50 Hz max.) is tuned near the 394.401 nm Al(I) resonant transition, expanded and collimated, and passed through the ablation plume.

To make an interferogram of the ablation plume, the dye laser beam is first split into two beams by an uncoated, quartz interferometry flat turned 70° from normal incidence. Due to internal reflections and the dimensions of the quartz flat, there are actually four beams leaving the flat. The primary transmitted beam and the primary reflected beam are utilized for the ablation photography in this work and the ratio of their energies is approximately 4:1. The primary transmitted beam is enlarged 28X with a Galilean telescope (FL=17 mm & FL=480 mm) and



Fig. 3.1(a) Experimental configuration of the main optics table. I=Iris, L=Convex Lens, M=Mirror, RM=Rotatable Mirror, and BS=Beam Splitter.



Fig. 3.1(b) Experimental configuration of the main optics table, continued. A silicon photo-diode at (A) measures the timing of the KrF laser and a p-i-n diode at (B) measures the timing of the dye laser. A thin, quartz flat splits part of the KrF off onto a calorimeter (8.8 V/J) at (C) in order to monitor the shot-to-shot energy of the Lumonics laser. The calorimeter is calibrated after an experiment using a second calorimeter (0.39 V/J) at (D) and varying the impinging KrF laser energy by placing thin quartz flats in the beam at (E). The KrF laser energy is also monitored during an experimental run by placing the second calorimeter at (F) both before and after the exposure of each holographic plate.

used as the object beam in the interferometer, passing through the ablation plume in the vacuum chamber. The primary reflected beam is enlarged 16X with a Galilean telescope (FL=17 mm & FL=280 mm) and used as the reference beam. Both beams have an iris located at the focal point in the telescope to remove uncollimated laser light from the photographing process. Mirrors are oriented to direct the collimated beams onto the holographic plate (Agfa-Gavert 8E56, sensitive to 350 nm - 560 nm light) and to create approximately equal traveling distances between the beam splitter and the holographic plate. In order for an interference pattern to form at the holographic plate, it is necessary that the difference in these pathlengths be less than coherence length of the laser beam, which is estimated to be about 5 cm. A 5.1x5.1 cm blue glass filter with peak transmission at 400 nm is mounted on the 4.4x4.4 cm entrance window of the holographic camera to keep out excess room light during the experiment.

The experimental procedure is as follows. First, the dye laser is pulsed without the Lumonics excimer laser firing (thus no ablation plume) to expose the holographic plate with a without-the-plume interference pattern. Next, the rotatable mirror in the reference beam is turned very slightly to move the reference image ≈ 1 mm on the holographic plate. Then the KrF and dye lasers are pulsed to form a second interference pattern on the holographic plate, this pattern containing the ablation plume information. Four such double-pulsed holographic interferograms are taken on each 4x5 inch holographic plate before the holographic camera is reloaded.

Careful attention is given to the temperature of the dye laser grating during an experimental run. The wavelength output calibration (the dye laser control unit readout versus the actual output wavelength) changes 0.0006 nm/°C, which is 20% of the 0.0031 nm FWHM laser intensity bandwidth, so the grating temperature and the room temperature were monitored to the nearest 0.1°C. For the resonant holograms presented here, the dye laser output was recalibrated (see Chapter 4) if the grating temperature changed by \approx 1°C from the time of the initial output calibration.

For Dye-Laser-Resonance-Absorption-Photography (DLRAP), the holographic camera light-tight-box is removed from the camera holder and the object beam then passes through the blue glass plate, neutral density filters (not shown), and into the DLRAP camera, which uses Polaroid Type 47 film.

Figure 3.1(b) shows the location of the p-i-n photo-diodes and the calorimeters on the optics table. The UV enhanced silicon photo-diode for the Lumonics, (A), and the p-i-n photo-diode for the dye laser, (B), are permanently fixed to the optical table. Their rise-times to the laser pulses are on the same order as the length of the laser pulses, thus allowing the time difference between the peaks of the laser pulses to be measured on an oscilloscope. (For a description of the timing circuitry and an example of the timing signals, see Appendix 4.)

During the experiment, the Lumonics laser energy decreases with time and use, dropping nearly 50% after 30 - 45 minutes of intermediate use. Usually, 33% of the Lumonics KrF gas fill was changed after the exposure of each holographic plate; i.e., about every 10 - 20 minutes. Therefore, the Lumonics energy was usually allowed to drop $\approx 25\%$ during the course of exposing the four, double-pulse holograms for each holographic plate. To monitor the shot-to-shot Lumonics laser energy, a

portion of the beam is split off with a thin quartz flat onto a sensitive calorimeter (8.8 V/J), located at position (C). The peak of the calorimeter signal is then proportional to the total energy impinging onto the aluminum target.

To calibrate calorimeter (C) after each experiment, the target and the back flange of the vacuum chamber were removed from the Lumonics beam path and a second, large-area calorimeter (0.39 V/J) was placed behind the vacuum chamber at position (D). The signals on both calorimeters were recorded as a function of the number of thin, quartz flats placed in the beam path at position (E), and thus as a function of laser energy. (For a description of the calorimetry circuitry, see Appendix 4. For an example of the calorimetry signals and of a calibration curve, see Appendix 5.) To provide a day-to-day monitoring of the Lumonics energy changes, the second calorimeter (0.39 V/J) was placed at position (F) both before and after the exposure of each holographic plate.

The Lumonics Laser Spot Shape

Figure 3.2 shows a photograph of a selection of laser ablated, solid aluminum targets (99.8% pure). Spot #2 was the result of >300 laser pulses in October 1992 and was used for about half of the holographic data presented in this work. The laser intensity profile at that time was fairly uniform. Visually, the ablation spot size appears to grow with the number of incident ablation pulses. While the center of the beam etches the target noticeably on the first pulse, the etching due to the wings of the ablating laser takes more pulses to visually notice. (e.g., spot #1





Fig. 3.2 (a) is a photograph of three, 99.8% pure aluminum targets used in KrF laser ablation experiments and (b) labels each of the ablation spots. Spots #1 - #2 were taken in October 1992 and are representative of laser spots from the period July through December, 1992. Spot #2 was etched with more than 300 ablation pulses and a spot size of 0.10 cm² was measured for it, slightly smaller than the visible area showing because the apparent spot size increases with the number of ablation pulses. Spots #3 - #6 were taken in March 1993 and are representative of laser spots from the period January through June, 1993. Note the double hot spot in the KrF laser intensity profile. Spot #3 was etched with about 50 ablation pulses and a spot size of 0.065 cm² was measured for it. Spots #7 - #9 were taken in June 1993 with the aluminum target inclined 69 ° to the KrF laser beam. Spot #9 was etched with 55 ablation pulses and a spot size of 0.049 cm² was measured for it. appears smaller than spot #2). Therefore a spot size slightly smaller than the actual spot (#2) was used for the fluence calculations for the October 1992 data (0.10 cm^2).

In about January 1993, the laser intensity profile of the Lumonics laser changed and remained so through the rest of the data taken for this work. It exhibited two distinct hot spots of unequal intensity, as in spots #3-#6. Attempts to prevent one of the two spots from hitting the target while still obtaining a detectable fringe shift using RHI were unsuccessful, so the laser spot was used as is. Although none of the spots on the center target correspond to particular holograms shown in this work, it is the most clear example of ablated target spots for the data taken between January and June, 1993. Much of the holographic data taken in this time period is not presented in this work because there were often two, unequal plumes expanding away from the target; thus the holograms presented from that time period are the least affected by the double-spot laser intensity profile.

Again the appearance of the ablation spot changes with the number of pulses. Spot #5 is due to only several pulses and shows only the beginning of material removal from the surface. Spots #3 and #6 are due to more shots, and they show the double craters that form with that particular laser intensity profile. Spot #4, which was made with the most ablation pulses of the four spots, shows a ring of soot around the spot due to redeposition of ablated material. For the purpose of fluence calibrations, the size ablation spot was measured to be the size of a rectangle just encompassing the two laser spots (e.g., 0.065 cm² for spot #3 for \approx 50 ablation pulses). The Lumonics twin-spot was exaggerated for the deposition experiment because the target was tilted 69° to allow the Lumonics laser beam access to the target. The silicon wafer was placed 1.5 cm away from the ablation target and 55 ablation pulses formed spot #9, which was used for both the deposition holograms presented here as well as the forming of the aluminum thin film on the wafer. The wafer is <100> crystalline silicon coated with 115 nm of SiO₂ followed by 493 nm of amorphous silicon. The aluminum thin film was deposited on the amorphous silicon layer. Here, the spot size for the fluence calibration was determined to be the sum of the two separate ablation areas for spot #9 (0.049 cm²).

Developing And Reconstructing The Holograms

Once the holographic plates were exposed, they were developed in a darkroom using the procedure given in Table 3.1. They were then reconstructed with an expanded HeNe laser using the setup shown in

Table 3.1 Photographic developing process for the holographic plates.

1)	3 - 6 min.	D-19 Photographic Developer
		(visually verify proper development using
		darkroom lighting)
2)	30 sec.	Stop Bath
3)	4 min.	Rapid Fixer
4)	10 sec.	Water (continue using regular
,		lighting)
5)	4 min.	Hypoclear
6)	30 sec.	Water
7)	30 sec.	Photoflo



Figure 3.3 Schematic For Reconstructing The Hologram (see Fig. 2.2). HP=Holographic Plate, C= Camera, E=Light Tight Enclosure, and A-S=Aperture and Shutter.

Fig. 3.3. The holographic plate orientation with respect to the laser is reversed from the direction of exposure to utilize a more intense reconstructed image of the plume. Reconstruction with a laser wavelength other than the exposure wavelength does not change the image (for a transmission hologram), just the image location and magnification (Ves79). The reconstructed image was focused through a camera shutter (which doubled as an iris) and into a camera using Polaroid 667 film was placed in the path of the reconstructed image. The shutter speed was adjusted to vary the exposure (and thus the contrast) of the image on the film.

Plume Expansion In RF/Plasma Environments

Figure 3.4 shows the experimental setup to produce a steady-state (for < 10 sec) RF-plasma in the region of plume expansion in order to

observe the effects of an RF-plasma on plume expansion. Since this is the focal point of Chi Hong Ching's forthcoming dissertation, Fig. 3.4 is shown here as an example of an application of RHI and its explanation will be brief.

A 7 MHz, \approx 50 W, RF generator was connected to a pair of 8.4 cm diameter, brass, parallel plates separated by 4.8 cm inside the the vacuum chamber. The aluminum target face was positioned so that it was just slightly outside the parallel plates. This allowed the plume to expand between the parallel plates while not letting the target significantly perturb the electric field between the plates.

The turbopump port valve was closed 99% and argon gas was leaked onto the vacuum chamber through a leak valve (not shown) to obtain a constant pressure of 1 Torr. The RF voltage was turned on for about 3 seconds before the ablating laser was fired, thus allowing the RF circuitry and the plasma sufficient time to equilibrate.



Fig. 3.4 Schematic of the RF-Argon-Plasma setup. The brass, parallel plates have a diameter of 8.4 cm and are separated by 4.8 cm. The plasma current is measured using the Pearson transformer.

CHAPTER 4 EXPERIMENTAL RESULTS AND DISCUSSION

Determining λ_0

To determine the center of the atomic resonance with respect to the dye laser wavelength control unit, a piece of paper is placed in the object beam path while the wavelength is scanned near the resonance center during KrF laser ablation of an aluminum target. The absorption image of the plume on the paper is observed to determine the wavelength of most absorption and then four holograms (one holographic plate) are taken across the anticipated line center using the smallest wavelength increment available on the grating control unit (0.001 nm - 0.002 nm). The reference beam mirror is turned in the same direction for these four holograms for ease of interpretation.

Figure 4.1 shows four sample holograms in a 1.0 Torr argon background, at 2.5 μ s, and with an ablation laser fluence of ~2.3 J/cm². First of all, holograms 4.1(b) and 4.1(c) show fringe shift toward the target while hologram 4.1(d) shows fringe shift away from the target. This confirms the change in the index of refraction from n<1 to n>1 when crossing from one side to the other side of the line center (see Fig 2.3). (The polarity of the change will be experimentally determined in the next section.) Secondly, hologram 4.1(c) has the greatest amount of absorption. Therefore, hologram 4.1(c) is chosen as $\Delta\lambda = 0.000$ nm for



Fig. 4.1 Calibrating the dye laser to the aluminum line at 394.401 nm. (a) Photographic orientation; holograms (b), (c), (d) and (e) were taken in 1 Torr Argon gas at (2.50 μ s, =2.3 J/cm²). (b), (c) and (d) were taken successively with a wavelength increment of 0.001 nm. Since (c) has the most absorption and the fringe shift direction changes between (c) and (d), hologram (c) was chosen to be line center, $\Delta\lambda = +0.000$ nm. Holograms (b) and (d) thus have calibrated wavelengths of -0.001 nm and +0.001 nm, respectively. Hologram (e) shows no fringe shift at $\Delta\lambda = +0.493$ nm and thus all fringe shift in (b), (c) and (d) is due to resonant effects. (From laser spot #2 of Fig. 3.2(b).)

the purposes of interpreting holograms within this data run. Hologram 4.1(e) has a large $\Delta\lambda$, +0.493 nm, and no fringe shift, thus showing that the fringe shift seen in holograms 4.1(b)-4.1(d) are due only to resonant effects.

For better fringe shift interpretation, it was chosen to turn the reference beam mirror in such a direction as to have the fringes shift away from the target. Therefore, for $\Delta\lambda < 0$ the mirror was turned in one direction while for $\Delta\lambda > 0$ the mirror was turned in the other direction.

Resonant Refraction Effects

Figure 4.2 shows the resonant refraction effects on holograms for operating on both sides of the line center. These effects are important because, while some light is passing through the ablation plume unrefracted, the light passing through the more dense regions of the plume may be refracted. The refracted and nonrefracted light can then interact and alter the fringe appearance of the holograms.

Figure 4.2(a) shows that for n>1, the laser ablation plume acts similar to a convex glass lens and focuses the light toward the center of the plume. Thus, if the most dense region of the ablation plume has separated from the target surface, the edges of the plume will appear unaltered while the center of the plume will appear to be a confusing pattern of black and white, as in Fig. 4.2(c).

Figure 4.2(b) shows that for n<1, the laser ablation plume acts similar to a concave glass lens (or a convex plasma) and defocuses the light away the center of the plume. Thus, if the most dense region of the ablation plume has separated from the target surface, the center of the



Fig. 4.2 (a) and (b) show refraction effects of the dye laser beam encountering an ablation plume with (n > 1) and (n < 1), respectively (Ves79). Refraction effects are enhanced near resonance as evidenced by holograms (c) and (d), taken in vacuum at 0.50 µs and ≈ 2.4 J/cm². Hologram (c) was taken with +0.004 nm and exhibits crisscrossing fringes at its center, thus the ablation plume had (n > 1). Hologram (d) was taken with -0.003 nm and exhibits non-fringe lines encircling the plume and light located on the target in the hologram, thus the ablation plume had (n < 1). (From laser spot #2 of Fig. 3.2(b).)
plume will appear unaltered (or dimly lit) while the edges of the plume will appear to be altered, as in Fig. 4.2(d). Note the light that appears on the target in the hologram at the left and right sides of the plume which can only come from refraction away from the plume center. Given the calibrated wavelengths of Figs. 4.2(c) and (d) are +0.004 nm and -0.003 nm, respectively, the polarity of the phase shift response curve is then experimentally confirmed to be

$$\operatorname{sgn}\{\operatorname{Re}[\tilde{n}-1]\} = -\operatorname{sgn}\{\omega - \omega_{o}\}$$
(4.1)

in agreement with Eqn. 2.25 and as shown in Fig. 2.3.

For the purposes of better fringe shift interpretation, a large enough $|\Delta\lambda|$ was chosen to attempt to minimize resonant refraction effects while retaining the resonant particle sensitivity of smaller $|\Delta\lambda|$'s. Also, it is possible that the resonant refraction effects could be decreased by properly imaging the ablation plume during the reconstruction process, assuming that the refracting region in the ablation plume was small (i.e., on the order of the depth of focus for the imaging setup).

Fringe Shift Interpretation

Figure 4.3 shows the basic method for interpreting holographic interferometry fringe shifts. The holograms are enlarged 200% using a photocopier and the fringes are traced by hand onto a clean sheet of paper placed on a light board. Non-fringe features of the hologram, such as refraction, diffraction, absorption and uninterpretable or distracting features, are thereby removed from the working image. In the region of the ablation plume, the unshifted fringes are also drawn onto the fringe tracing, as shown by the dashed lines in Fig. 4.3(a).



Fig. 4.3 (a) shows the method of fringe shift interpretation (Ves79). The points of incremental fringe shift are determined along the unshifted, dotted lines by their intersection with other shifted fringes and then, not shown here, the equi-fringe shift points are connected with a smooth line to form equi-fringe shift contours. Hologram (b) was taken in 1 Torr Argon gas at (4.00 μ s, -0.006nm, =2.4 J/cm²). (c) is an equicontour plot of (b) where the dashed line is a shock wave and the solid lines correspond to increasing fringe shifts; from outer to inner line: (0.5 FS =7.0x10¹³ cm⁻²),(1.0 FS=1.4x10¹⁴ cm⁻²), (1.5 FS=2.1x10¹⁴ cm⁻²), (2.0 FS =2.9x10¹⁴ cm⁻²), (2.5 FS=3.6x10¹⁴ cm⁻²), (3.0 FS=4.4x10¹⁴ cm⁻²), (3.5 FS=5.2x10¹⁴ cm⁻²), and (4.0 FS=6.0x10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)

Next, the photocopy of the hologram is removed from the light board and a new sheet of paper is placed over the first tracing. The points of equal phase shift are determined and recorded on the new sheet, and then these points are connected with a smooth curve. In Fig. 4.3(a) for example, all of the points labeled "2" are recorded on the overlaying sheet of paper and then connected with an oval passing though each of the six points. To determine half-fringe shifts, the points where the fringe shifts halfway between two reference fringes (dashed lines) is used.

Figure 4.3(b) and (c) show an example of a hologram at 4.00 μ s in 1 Torr argon gas, and its equi-line-density-contour plot. The dashed line represents the shock wave and the solid lines represent incremental 0.5 fringe shifts, to a maximum of 4.0 fringe shifts at the center of the plume, which corresponds to a maximum aluminum neutral line-density of 6.0×10^{14} cm⁻² for this hologram. Again, all of the aluminum neutral line-densities presented in this work assume a constant kinetic temperature of 0.3 eV in a Maxwellian ablation plume in quasi-equilibrium.

Unfortunately, there are several points where uncertainties are introduced in the interpretation process. First of all, uncertainties are introduced in the process of interpreting the black and white pattern on the hologram to determine what are fringes versus what are refraction and diffraction effects. Areas near the center of ablation plumes sometimes have a certain amount of absorption or resonant refraction effects that brought uncertainty into the location of the shifted fringes, including what would appear to be additional fringe lines in the center of the ablation plume. Either these areas were not interpreted at all or human discernment was used to interpret the recorded information. Secondly, uncertainties are introduced in the process of choosing the location of the unshifted reference fringes (dashed lines). The reference fringes are actually slightly curved due to the dye laser beam wavefront not being of constant phase over the entire expanded beam, so an uncertainty of up to 0.25 fringe shift is introduced in this step. Third, uncertainties are introduced in the process of accurately tracing the points of equal fringe shift and, finally, in the process of interpretively drawing a smooth curve between the points. Digital fringe shift recording systems (such as CCD cameras) may be able to reduce uncertainty in the "tracing" process but complex fringe shift interpretation programs are needed to distinguish between fringe shift, refraction, and diffraction. For this reason, the holograms in this work were interpreted by tracing.

Shot-To-Shot Consistency

Figure 4.4 shows a set of ten holograms taken in vacuum at the same time, $0.51 - 0.53 \ \mu$ s, with approximately the same ablating KrF laser fluence, 2.2 - 2.6 J/cm², and with the same wavelength, $\Delta \lambda = -0.013$ nm. This set of tracings shows that there is a wide range of variability in the shape of the ablation plume even when the experimental conditions are supposedly the same. Fig. 4.4 also demonstrates that the size of the ablation plume is somewhat dependent on fluence. The tracings are in the order of fluence and are, in general, larger for higher fluences. The differences between the holograms are most likely due to the changing laser intensity profile from shot-to-shot and to some degree the changes in the ablation surface caused by previous ablation pulses. It is also possible that the plumes had different temperatures, which would



Fig. 4.4 Ten holograms taken at -0.013 nm in vacuum to demonstrate shot-to-shot consistency. Lines correspond to increasing fringe shifts; from outer to inner: $(0.5 \text{ FS}=1.6 \times 10^{14} \text{ cm}^{-2})$, $(1.0 \text{ FS}=3.2 \times 10^{14} \text{ cm}^{-2})$, $(1.5 \text{ FS}=4.8 \times 10^{14} \text{ cm}^{-2})$, and $(2.0 \text{ FS}=6.4 \times 10^{14} \text{ cm}^{-2})$. The exception is (a) where both solid lines correspond to 0.5 FS and the line-density is nearly zero in the center of the ablation plume. (a) and (b) were at $\approx 2.2 \text{ J/cm}^2$, (c), (d), (e), (f), (g) and (h) were at $\approx 2.3 \text{ J/cm}^2$, (i) was at $\approx 2.6 \text{ J/cm}^2$. Note that (b), (c) and (e) show possible effects of the double focal spot of the KrF laser. (From similar to laser spots #3-#6 of Fig. 3.2(b).)



Fig. 4.5 Line-density versus wavelength at 0.5 μ s for two fluences to show consistency of line-density measurement as a function of wavelength. (lower fluence: from spot #2 of Fig. 3.2(b); higher fluence: from similar to spots #1 - #2 of Fig3.2(b))

produce some variation in the observed fringe shift. The variations shown in Fig. 4.4 mean that hologram interpretation of a set of holograms where one parameter is varied can only be approximate at best.

Two series of holograms were taken at 0.5 μ s at various wavelengths in order to determine the consistency of measuring the linedensity for an ablation plume over a range of dye laser wavelengths. The first set of holograms used a fluence of 2.0 - 2.7 J/cm² and the second set used a fluence of 4.2 - 4.7 J/cm². Figure 4.5 shows the results as a function of wavelength for the two scans at a point 0.7 cm away from the target along the line of symmetry of the ablation plume.

First of all, note that absorption prevented the measurement of fringe shift at 0.7 cm on the hologram for $|\Delta\lambda| < 0.009$ nm for the higher

laser fluence case. Secondly, note that the measured aluminum neutral line density seemingly increases with increasing $|\Delta\lambda|$. This is because the measured fringe shifts at this point on each of the holograms were around 1.0 fringe shift and were measured to the nearest 0.25 fringe shift. Thus, Fig. 4.5 looks similar to a theoretical curve showing the aluminum neutral line-density interpretation as a function of wavelength for a constant 1.0 fringe shift. Third, note that the ratio of the line-density averages for the two fluences is ≈ 2.5 , which is comparable to, yet larger than, the ratio of the average laser fluences, ≈ 1.9 . This is consistent in principle with the findings of Dreyfus (Dre91), where the LIF signal (and thus the particle density) for copper ablation was experimentally determined to be proportional to $\exp(-1/F)$ for fluences identical to this work, where F is the laser fluence, although Dreyfus measured a gain of $\approx 25X$ in the LIF signal for the same fluence increase as presented in Fig. 4.5.

It is suggested that if the fringe shifts at an identical point on a series of holograms were measured with greater certainty (such as 5.00 ± 0.25 fringe shifts instead of 1.00 ± 0.25 fringe shifts), then the graph of the measured line density versus wavelength would be more constant as a function of wavelength for each fluence. It is also entirely possible that the temperatures in the plumes at the two different fluence ranges could be different, and thus affecting the outcome of Fig. 4.5. A temperature of 0.3 eV was assumed for both fluence ranges since the individual plume temperatures could not be determined at that time.

Laser Ablation In Vacuum And Low Pressure, Inert, Background Gas

Resonant holographic interferometry time scans of laser ablation plume expansion, within the fluence range of $2 - 6 \text{ J/cm}^2$, were taken in vacuum (0.2 mTorr air) and in argon gas backgrounds for the pressures 14 mTorr, 52 mTorr, 210 mTorr, 1 Torr, and 35 Torr. In addition, plume expansion in 1 Torr argon plasma environment was explored. The plume expansions as a function of time were very similar for the background gas pressures up to 210 mTorr, so their plume expansion interpretations will be handled together and compared to the laser ablation plume expansion-in-vacuum theory used by Utterback, et al. (Utt76), Talents (Tal83), and Kools, et al. (Koo92). Then the 1 Torr and 35 Torr data will be presented and, finally, the summaries of the data will be presented.

Figure 4.6 shows a time scan in a series of holograms in vacuum. The ablation plume begins close to the surface and expands away from the target at later times. It is assumed that the aluminum neutral atoms experience collisions and that almost all of the electrons have recombined with the atoms at very early times ($< \approx 0.1 \ \mu$ s). The peak aluminum neutral line-density on each hologram (> 5x10¹⁴ cm⁻²) is located at the surface at early times (0.29 μ s and 0.50 μ s) and then separates from the surface at later times (0.70 μ s, 0.90 μ s, and 1.11 μ s). This corresponds to inferred particle densities of up to 1x10¹⁵ cm⁻³. Also, the peak aluminum neutral line density decreases at later times to $\approx 1x10^{14}$ cm⁻² (at 1.11 μ s).



Fig. 4.6 Time scan in vacuum. Hologram (a) taken at $(0.29 \ \mu s, +0.004 \text{nm}, \approx 2.4 \text{ J/cm}^2)$ and hologram (b) taken at $(0.50 \ \mu s, +0.004 \text{nm}, \approx 2.4 \text{ J/cm}^2)$. (c) is an equicontour plot of (a), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS} = 4.3 \times 10^{13} \text{ cm}^{-2})$, $(1.0 \ \text{FS} = 9.0 \times 10^{13} \text{ cm}^{-2})$, $(2.0 \ \text{FS} = 1.9 \times 10^{14} \text{ cm}^{-2})$, and $(3.0 \ \text{FS} = 3.0 \times 10^{14} \text{ cm}^{-2})$; (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS} = 4.3 \times 10^{13} \text{ cm}^{-2})$, $(1.0 \ \text{FS} = 9.0 \times 10^{13} \text{ cm}^{-2})$, $(2.0 \ \text{FS} = 1.9 \times 10^{14} \text{ cm}^{-2})$, and $(3.0 \ \text{FS} = 3.0 \times 10^{14} \text{ cm}^{-2})$; (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS} = 4.3 \times 10^{13} \text{ cm}^{-2})$, $(1.0 \ \text{FS} = 9.0 \times 10^{13} \text{ cm}^{-2})$, $(2.0 \ \text{FS} = 1.9 \times 10^{14} \text{ cm}^{-2})$, $(3.0 \ \text{FS} = 3.0 \times 10^{14} \text{ cm}^{-2})$, $(4.0 \ \text{FS} = 4.2 \times 10^{14} \text{ cm}^{-2})$, and $(5.0 \ \text{FS} = 5.5 \times 10^{14} \text{ cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.6 Time scan in vacuum, continued. Hologram (e) taken at $(0.70 \ \mu\text{s}, +0.004 \text{nm}, \approx 2.4 \text{ J/cm}^2)$ and hologram (f) taken at $(0.90 \ \mu\text{s}, +0.004 \text{ nm}, \approx 2.4 \text{ J/cm}^2)$. (g) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS} = 4.3 \times 10^{13} \text{ cm}^{-2})$, $(1.0 \text{ FS} = 9.0 \times 10^{13} \text{ cm}^{-2})$, $(1.5 \text{ FS} = 1.4 \times 10^{14} \text{ cm}^{-2})$, $(2.0 \text{ FS} = 1.9 \times 10^{14} \text{ cm}^{-2})$, and $(2.5 \text{ FS} = 2.5 \times 10^{14} \text{ cm}^{-2})$; (h) is an equicontour plot of (f), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS} = 4.3 \times 10^{14} \text{ cm}^{-2})$; (h) is an equicontour plot of (f), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS} = 4.3 \times 10^{13} \text{ cm}^{-2})$, $(1.0 \text{ FS} = 9.0 \times 10^{13} \text{ cm}^{-2})$, $(1.5 \text{ FS} = 1.4 \times 10^{14} \text{ cm}^{-2})$, and $(2.0 \text{ FS} = 1.9 \times 10^{14} \text{ cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.6 Time scan in vacuum, continued. Hologram (i) taken at (1.11 μ s, +0.004 nm, =2.4 J/cm²). (j) is an equicontour plot of (i), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS=4.3×10¹³ cm⁻²), (1.0 FS=9.0×10¹³ cm⁻²), and (1.5 FS =1.4×10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.7(a) Total number of alumunim neutral atoms, N_o , versus time in vacuum.



in vacuum.

Figure 4.7(a) shows the total number of aluminum neutral particles on the hologram, N_0 , as a function of time for the vacuum time scan tracings presented in Fig. 4.6. The parameter N_0 is determined by tracing the holograms onto fine graph paper and counting the squares in each region. For example, the number of squares between 0.5 fringe shift and 1.0 fringe shift is counted, converted into an area in cm², and multiplied by the calculated aluminum neutral line-density for the 0.5 fringe shift equicontour aluminum neutral line to obtain an estimate of the total number of particles represented between the 0.5 and 1.0 fringe shifts. This process is repeated for each section of the hologram and summed to approximate N₀. The value N₀ determined for the holograms at 0.70 μ s, 0.90 μ s, and 1.11 μ s is essentially the same, 3x10¹³, and shows consistency with time even when the peak line density decreases from 2.5x10¹⁴ cm⁻² to 1.4x10¹⁴ cm⁻².

Figure 4.7(b) plots the farthest extent of the 1×10^{14} cm⁻² equicontour line as a function of time for the holograms in Fig. 4.6. The expansion velocity, v, was determined by calculating the best fit to a line passing through the first four points as well as the origin. The expansion velocity of 1.41 cm/µs is consistent with other recent measurements of the TOF most probable expansion velocity (Geo89, Wan91, Zhe89) and the leading edge expansion velocity (Dre91, Ven92a, Wal86) of laser ablation expansion in vacuum.

Figure 4.8 (a) through (1) shows another series of holograms in vacuum (0.2 mT air), except this time for a laser ablative deposition configuration in which the laser is incident on the aluminum target with an angle of $\approx 69^{\circ}$ and an amorphous-silicon-coated silicon wafer has been placed 1.5 cm away from the aluminum target. The expanded dye laser



Fig. 4.8 Deposition time scan in vacuum. Hologram (a) taken at $(0.25 \,\mu\text{s}, -0.016 \,\text{nm}, \approx 5.4 \,\text{J/cm}^2)$ and hologram (b) taken at $(0.52 \mu\text{s}, -0.009 \,\text{nm}, \approx 6.2 \,\text{J/cm}^2)$. An amorphous silicon wafer at room temperature is mounted 1.5 cm away from the aluminum target and is on the right side in the above holograms. (c) is an equicontour plot of (a), line corresponds to $(0.5 \,\text{FS}=2.0 \times 10^{14} \,\text{cm}^{-2})$; (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \,\text{FS}=1.1 \times 10^{14} \,\text{cm}^{-2})$ and $(1.0 \,\text{FS}=2.2 \times 10^{14} \,\text{cm}^{-2})$. (From laser spot #9 of Fig. 3.2(b).)



Fig. 4.8 Deposition time scan in vacuum, continued. Hologram (e) taken at (1.01 μ s, -0.009 nm, =5.6 J/cm²) and hologram (f) taken at (1.21 μ s, -0.016 nm, =6.2 J/cm²). An amorphous silicon wafer at room temperature is mounted 1.5 cm away from the aluminum target and is on the right side in the above holograms. (g) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS=1.1x10¹⁴ cm⁻²), (1.0 FS=2.2x10¹⁴ cm⁻²), (1.5 FS =3.3x10¹⁴ cm⁻²), (2.0 FS=4.4x10¹⁴ cm⁻²), (2.5 FS=5.5x10¹⁴ cm⁻²), and (3.0 FS =and 6.6x10¹⁴ cm⁻²); (h) is an equicontour plot of (f), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS=2.0x10¹⁴ cm⁻²), (1.0 FS=4.0x10¹⁴ cm⁻²), and (1.5 FS =5.9x10¹⁴ cm⁻²). (From laser spot #9 in Fig. 3.2(b).)



Fig. 4.8 Deposition time scan in vacuum, continued. Hologram (i) taken at $(1.69\mu s, -0.009 \text{ nm}, \approx 6.2 \text{ J/cm}^2)$ and hologram (j) taken at $(2.03 \ \mu s, -0.003 \text{ nm}, \approx 6.2 \text{ J/cm}^2)$. An amorphous silicon wafer at room temperature is mounted 1.5 cm away from the aluminum target and is on the right side in the above holograms. (k) is an equicontour plot of (i), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS}=1.1\times10^{14} \text{ cm}^{-2})$, $(1.0 \text{ FS}=2.2\times10^{14} \text{ cm}^{-2})$, $(1.5 \text{ FS}=3.3\times10^{14} \text{ cm}^{-2})$, $(2.0 \text{ FS}=4.4\times10^{14} \text{ cm}^{-2})$, and $(2.5 \text{ FS}=5.5\times10^{14} \text{ cm}^{-2})$; (l) is an equicontour plot of (j), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS}=4.4\times10^{14} \text{ cm}^{-2})$, and $(2.5 \text{ FS}=5.5\times10^{14} \text{ cm}^{-2})$; (l) is an equicontour plot of (j), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \text{ FS}=3.2\times10^{13} \text{ cm}^{-2})$ and $(1.0 \text{ FS}=6.9\times10^{13} \text{ cm}^{-2})$. (From laser spot #9 of Fig. 3.2(b).)



Fig. 4.8 Deposition time scan in vacuum, continued. Dye Laser Resonant Absorption Photographs (DLRAP's) that correspond to, but are not simultaneous to, holograms (a), (b), (e), (f), (i) and (j), plus two additional DLRAP's at later times: [(m), 0.21 μ s, -0.016 nm, \approx 5.4 J/cm²], [(n), 0.50 μ s, -0.009 nm, \approx 5.9 J/cm²], [(o), 1.00 μ s, -0.009 nm, \approx 5.8 J/cm²], [(p), 1.21 μ s, -0.009 nm, \approx 5.4 J/cm²], [(q), 1.61 μ s, -0.009 nm, \approx 5.4 J/cm²], [(r), 2.00 μ s, -0.003 nm, \approx 5.4 J/cm²], [(s), 4.02 μ s, -0.003 nm, \approx 5.8 J/cm²], and [(t), 5.02 μ s, -0.003 nm, \approx 4.8 J/cm²]. Note the transparent ruler to the right side of the silicon wafer and the wafer holder, and that the DLRAP image is inverted with respect to the holograms. (From laser spot #9 in Fig. 3.2(b).)

beam was not perfectly parallel to the target and wafer surfaces for this experimental run, so the separation distance appears to be only ≈ 1.25 cm.

The ablation laser energy was higher for this run, so the fluence is higher for Fig. 4.8 than for Fig. 4.6, even though the target is tilted and the spot size is larger. The ablation plume begins close to the surface and expands toward the silicon wafer at earlier times (0.25 μ s, 0.52 μ s, and 1.00 μ s). By 1.21 μ s, the ablation plume makes contact with the wafer and then by 1.69 μ s, the plume is observed to begin receding back to the aluminum surface. This does not mean that the aluminum neutrals flow toward the wafer and then recede back as an ocean wave advances and recedes on a beach, rather the aluminum neutrals are continually expanding outward and the RHI-detectable portion of the aluminum neutral atoms appears to extend from and recede back to the aluminum target surface. The peak line-density (>6x10¹⁴ cm⁻²) appears to remain at the target surface, although the holograms are greatly obscured by absorption until 1.69 μ s. The inferred particle density is estimated to be >5x10¹⁴ cm⁻³.

Figure 4.8 (m) through (t) shows a series of Dye-Laser-Resonance-Absorption-Photographs (DLRAP's) that correspond to the holograms in Fig. 4.8. During the experimental run, four holograms were taken and then the corresponding four DLRAP's were taken under the same conditions. Here also, the ablation plume is seen to expand toward the wafer at earlier times (0.21 μ s), make contact with the wafer at intermediate times (0.50 μ s, 1.00 μ s, 1.21 μ s, and 1.61 μ s), and recede from the wafer at later times (2.00 μ s, and 4.02 μ s). By 5.02 μ s, the ablation plume is no longer observable. Note that DLRAP is more sensitive than RHI, and thus the presence of the ablation plume was detected as late as $4.02 \mu s$ with DLRAP.

The total number of aluminum neutral atoms at 1.69 µs was 8.3×10^{13} cm⁻²; the parameter N_o could only be determined for one hologram (Fig. 4.8(i)) due to the absorption observed in this time scan. Since this hologram, at 1.69 μ s, was taken after the plume appeared to begin to recede back to the target surface, its measurement of No is perhaps four times lower than the true number of aluminum neutral particles leaving the target surface (as estimated from Fig. 4.8(g), which makes $\approx 3 \times 10^{14}$ aluminum atoms total). When compared with the first time scan in vacuum, the plume expansion in Fig. 4.8 is much more directional, and thus a greater number of particles (> 3 times) expanding in a more forward direction produce absorption for a greater distance away from the target, even for larger $|\Delta\lambda|$'s. The variation in plume expansion shape has been observed to depend upon experimental parameters such as fluence magnitude, fluence uniformity, background pressure, and the distance from the laser focal lens to the target (Sap92, Lin93). Therefore, the plume shapes presented here vary from run-torun as well as from shot-to-shot, as shown in Fig. 4.4.

Figure 4.9 shows the expansion of the 1×10^{14} cm⁻² equicontour line as a function of time for the vacuum-deposition holograms. The leading edge expansion velocity of ≈ 1.05 cm/µs is about the same as the first time scan in vacuum, ≈ 1.41 cm/µs.

Figure 4.10 shows a photograph of the amorphous-silicon-coated silicon wafer with the deposited aluminum thin film after 55 ablation pulses on spot #9 of Fig. 3.2(b), and Figure 4.11 shows a rendition of the deposition spot with an overlay of the laser ablation spot. Note the



Fig. 4.9 Expansion of the 1×10^{14} cm⁻² equicontour line in vacuum for the deposition case.

photographic orientations: up in Fig. 4.11 is to the right in Fig. 4.10, upward in the DLRAP's of Fig. 4.8 and downward in the holograms of Fig. 4.8. Figure 4.11 shows clearly that the ablation plume expanded differentially in the x- and y-direction (z defined as perpendicular to the target and substrate surfaces), and that it expanded more quickly in the direction corresponding to the narrow width of the ablation spot. This sideways motion has been previously observed in atmospheric pressure air and modeled using hydrodynamic equations (Mio92); i.e., the expansion of the ablation plume particles in the sideways direction is governed primarily by self-collision and thus self-diffusion mechanisms $\underline{\Gamma} = -D\nabla n$, especially for expansion into vacuum and lower pressure background gases. For an initially (<0.1 µs) uniform particle density, the more narrow dimension of the laser spot corresponds to a higher density gradient and thus a higher expansion velocity. Thus a circular ablation



Fig. 4.10 Laser ablation deposition substrate (amorphous silicon) after 55 KrF ablation laser pulses on a 99.8% pure aluminum target (spot #9 of Fig. 3.2(b)). Target-substrate separation distance was 1.5 cm; substrate was at room temperature.

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Fig. 4.11 The laser ablation spot from the target (spot #9 of Fig. 3.2(b), here shown as white) overlaying a depiction of the aluminum thin film deposited on the amorphous silicon substrate (from Fig. 4.10, here shown as various lines and shades) in the original orientation. Note that the direction of greater lateral plume expansion corresponds to the narrower dimension of the ablation spot.

spot would produce a circular thin film (in vacuum and lower background pressures) while a (double spot) rectangular ablation spot, as shown here, produces an oval thin film with the larger dimension turned 90° with respect to the longer ablation spot dimension (Mio92).

The shape of the thin film in Fig. 4.11 is therefore an indication that aluminum-aluminum collisions play a significant role in the expansion of the ablation plumes studied here. It is also a confirmation of the previously utilized assumption that the ablation plume can be considered to be Maxwellian in distribution from the collisions at early times, and in quasi-equilibrium over the duration of the dye laser pulse (20 ns). This is not to say that a significant percentage of ablated particles are experiencing collisions at the same time the hologram is taken (e.g., $0.2 - 2.0 \mu$ s), but rather the collisions the ablated particles experience at very early times, when the ablation plume has a high enough particle density, determine and "fix" the velocity distribution that is observed at later times when collisions are no longer significant.

It has been calculated using a Monte Carlo simulation that a quasi-Maxwelliam velocity distribution is obtained for a laser ablation plume if the average number of collisions per ablated particle is greater than ≈ 3 , which is based on the ablative removal of 0.8 monolayers per pulse for the specific case investigated (Noo87). It has therefore been suggested that it would be difficult to produce an ablation plume whose particles experimence less than 3 collisions on average and thus not have a quasi-Maxwellian velocity distribution (Kel88).

The width of the central, two thin film regions in Fig. 4.11 (inside the thick line), as viewed by the dye laser (left to right), is approximately the same width as the ablation plume in Fig. 4.8(e). The thickness of this region can be approximately calculated by assuming that most of the deposited particles are in this region, that the film has a diamond-basedpyramid shape, and by assuming that the total number of aluminum atoms is $\approx 3 \times 10^{14}$ (from above). Since the density of solid aluminum is 6.02×10^{22} cm⁻³ and since the thin film was made with 55 ablation pulses, the estimated maximum height of the thin-film-pyramid is calculated to be ≈ 35 nm, assuming 100% of the Al neutrals impinge upon the surface and stick. The wafer that was used in this experiment was crystalline silicon ($\approx 500 \ \mu m$ thick) over which a 115 nm coating of SiO₂ was deposited, followed by a 493 nm layer of amorphous silicon (Pen93). The aluminum thin film was deposited onto the amorphous silicon layer. Attempts to measure the aluminum thin film thickness using the four point resistance probe method (Pen93) were unsatisfactory. The thin film appeared to be non-conductive, perhaps indicating an oxide layer had formed on the thin film. Also, the morphology of the thin film was unknown, making it difficult to use a tabulated value of resistivity in order to calculate the thin film thickness.

A zeroeth order estimate of the target ablation rate can be made using the measured spot size of 0.049 cm² for the ablation spot on Fig. 4.11. The ablation rate of 3×10^{14} aluminum atoms per pulse corresponds to a surface etching rate of 1 nm/pulse, which is close to the etching rate of 5 - 10 nm/pulse measured by Kelly and Rothenburg for 2.3 J/cm² ablation of aluminum (Kel85). Again, nearly half of the ablated atoms may be redeposited on the target.

A zeroeth order estimate of the amount of energy absorbed by the target per ablated particle can be made. With up to 300 mJ reaching the target surface (for Fig. 4.8) and a reflectivity of 92.5%, the aluminum target absorbed 22.5 mJ total energy to thermally ablate (expel) $\approx 3 \times 10^{14}$ particles, which is ≈ 466 eV absorbed by the target per ablated particle. Since the kinetic energy of each particle is on the order of 1 eV, almost all of the energy absorbed by the target remains in the target and is thermally dissipated.

Figures 4.12, 4.14, and 4.16 show the holograms and tracings for time scans in 14 mTorr, 52 mTorr, and 210 mTorr argon background gas, while Figures 4.13, 4.15, and 4.17 show N₀ and the expansion of the $1x10^{14}$ cm⁻² equicontour for each of the holograms. As in the vacuum cases, the ablation plume begins near the target at earlier times and the peak line-density recedes back to the surface at later times. Likewise, N₀ and the expansion velocity are close to the values measured for the two vacuum cases. The only exception to this is for 52 mTorr argon gas, where in Fig. 4.14 (h) the minimum detectable fringe shift (0.5 FS) has completely separated itself from the target surface.

Figure 4.18 compiles the information about the five time scans presented thus far. The values of N₀ in Fig. 4.18(a) for the most part range from $\approx 1 \times 10^{13}$ aluminum neutral atoms to 1×10^{14} aluminum neutral atoms. These are actually low estimates of the number of ablated particles because: 1) the step method for approximating the number of particles between each of the fringe shift equicontour lines gives low approximations, 2) the earlier-time holograms which would better represent the total number of particles often had too much absorption to make counting possible, and 3) only one type of particle was counted. It is again suggested that the total number of aluminum atoms ablated from the target surface (as neutrals, ions, and molecules) is in the range of 1×10^{14} - 1×10^{15} for the range of fluences used here.

Figure 4.18(b) compiles the expansions of the 1×10^{14} cm⁻² equicontours for the five time scans. Their expansion velocities are relatively similar and have the range 1.05 cm/µs- 1.41 cm/µs. These expansion velocities can be used to calculate the initial temperature of the collisional ablation plume using the relation (Ven91)



Fig. 4.12 Time scan in 14 mTorr argon gas. Hologram (a) taken at $(0.40 \ \mu\text{s}, -0.006 \ \text{nm}, \approx 2.4 \ \text{J/cm}^2)$ and hologram (b) taken at $(1.00 \ \mu\text{s}, -0.006 \ \text{nm}, \approx 2.4 \ \text{J/cm}^2)$. (c) is an equicontour plot of (a), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS}=7.0 \times 10^{13} \ \text{cm}^{-2})$ and $(1.0 \ \text{FS}=1.4 \times 10^{14} \ \text{cm}^{-2})$; (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS}=7.0 \times 10^{13} \ \text{cm}^{-2})$, $(1.0 \ \text{FS}=1.4 \times 10^{14} \ \text{cm}^{-2})$, $(1.5 \ \text{FS}=2.1 \times 10^{14} \ \text{cm}^{-2})$, and $(2.0 \ \text{FS}=2.9 \times 10^{14} \ \text{cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.12 Time scan in 14 mTorr argon gas, continued. Hologram (e) was taken at $(2.00 \,\mu\text{s}, -0.006 \,\text{nm}, \approx 2.4 \,\text{J/cm}^2)$. (f) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \,\text{FS}=7.0 \times 10^{13} \,\text{cm}^{-2})$ and $(1.0 \,\text{FS}=1.4 \times 10^{14} \,\text{cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.13(a) Total number of aluminum neutral atoms, $N_{\rm o}$, versus time in 14 mTorr argon gas.



Fig. 4.13(b) Expansion of the 1×10^{14} cm⁻² equicontour line in 14 mTorr argon gas.



Fig. 4.14 Time scan in 52 mTorr argon gas. Hologram (a) taken at $(0.20 \ \mu\text{s}, +0.008 \ \text{nm}, \approx 3.9 \ \text{J/cm}^2)$ and hologram (b) taken at $(0.50 \ \mu\text{s}, +0.008 \ \text{nm}, \approx 3.7 \ \text{J/cm}^2)$. (c) is an equicontour plot of (a), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS} = 9.7 \times 10^{13} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, and $(1.5 \ \text{FS} = 2.9 \times 10^{14} \ \text{cm}^{-2})$; (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS} = 9.7 \times 10^{13} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, and $(1.5 \ \text{FS} = 9.7 \times 10^{13} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS} = 1.9 \times 10^{14} \ \text{cm}^{-2})$, and $(1.5 \ \text{FS} = 2.9 \times 10^{14} \ \text{cm}^{-2})$. (From laser spots similar to #3 - #6 of Fig. 3.2(b).)



Fig. 4.14 Time scan in 52 mTorr, continued. Hologram (e) taken at $(0.80 \ \mu\text{s}, +0.008 \ \text{nm}, \approx 3.6 \ \text{J/cm}^2)$ and hologram (f) taken at $(1.20 \ \mu\text{s}, +0.008 \ \text{nm}, \approx 3.5 \ \text{J/cm}^2)$. (g) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS}=9.7 \times 10^{13} \text{cm}^{-2})$, $(1.0 \ \text{FS}=1.9 \times 10^{14} \ \text{cm}^{-2})$, and $(1.5 \ \text{FS}=2.9 \times 10^{14} \ \text{cm}^{-2})$; (h) is an equicontour plot of (f), line corresponds to $(0.5 \ \text{FS}=9.7 \times 10^{13} \ \text{cm}^{-2})$. (From laser spots similar to #3 - #6 of Fig. 3.2(b).)



Fig. 4.15(a) Total number of aluminum neutral atoms, $\rm N_{o}$, versus time in 52 mTorr argon gas.





Fig. 4.16 Time scan in 210 mTorr argon gas. Hologram (a) taken at $(0.23 \ \mu\text{s}, -0.015 \ \text{nm}, \approx 4.1 \ \text{J/cm}^2$) and hologram (b) taken at $(0.40 \ \mu\text{s}, -0.015 \ \text{nm}, \approx 4.7 \ \text{J/cm}^2$). (c) is an equicontour plot of (a), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS = $1.9 \times 10^{14} \ \text{cm}^{-2}$) and (1.0 FS= $3.7 \times 10^{14} \ \text{cm}^{-2}$); (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS= $1.9 \times 10^{14} \ \text{cm}^{-2}$), (1.0 FS= $3.7 \times 10^{14} \ \text{cm}^{-2}$), and (1.5 FS= $5.6 \times 10^{14} \ \text{cm}^{-2}$). (From laser spots similar to #3 - #6 in Fig. 3.2(b).)



Fig. 4.16 Time scan in 210 mTorr, continued. Hologram (e) taken at $(0.51 \ \mu\text{s}, -0.015 \ \text{nm}, \approx 4.8 \ \text{J/cm}^2)$ and hologram (f) taken at $(0.74 \ \mu\text{s}, -0.015 \ \text{nm}, \approx 4.8 \ \text{J/cm}^2)$. (g) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \ \text{FS}=1.9 \times 10^{14} \ \text{cm}^{-2})$, $(1.0 \ \text{FS}=3.7 \times 10^{14} \ \text{cm}^{-2})$, $(1.5 \ \text{FS}=5.6 \times 10^{14} \ \text{cm}^{-2})$, and $(2.0 \ \text{FS}=7.4 \times 10^{14} \ \text{cm}^{-2})$; (h) is an equicontour plot of (f), line corresponds to $(0.5 \ \text{FS}=1.9 \times 10^{14} \ \text{cm}^{-2})$. (From laser spots similar to #3 - #6 of Fig. 3.2(b).)



Fig. 4.16 Time scan in 210 mTorr, continued. Hologram (i) taken at (1.24 μ s, -0.015nm, \approx 4.9 J/cm²). (j) is an equicontour plot of (i), line corresponds to (0.5 FS=1.9x10¹⁴ cm⁻²). (From laser spots similar to #3 - #6 of Fig. 3.2(b).)



Fig. 4.17(a) Total number of aluminum neutral atoms, N_0 , versus time in 210 mTorr argon gas.



Fig. 4.17(b) Expansion of the 1×10^{14} cm⁻² equicontour line in 210 mTorr argon gas.



Fig. 4.18(a) Total number of neutral aluminum atoms, N_0 , versus time in vacuum and lower pressure argon gas backgrounds.



Fig. 4.18(b) Expansion of the 1×10^{14} cm⁻² equicontour line in vacuum and lower pressure argon gas backgrounds.


Fig. 4.19 Expansion of the ablation plume center in vacuum and lower pressure argon gas backgrounds.

$$k_{\rm B}T = \frac{(\gamma - 1)^2}{2\gamma} * \frac{m \ u_{\rm max}^2}{2} , \qquad (4.2)$$

where $\gamma=1.667$ for an ideal, monoatomic gas and u_{max} is the maximum expansion velocity measured for the laser ablation plume after the plume has become non-collisional. The velocities 1.05 cm/µs and 1.41 cm/µs then relate to initial plume temperatures 2.1 eV and 3.7 eV, respectively. This falls within the temperature range measured by Dreyfus using LIF on copper ablation for similar fluences: ≈ 0.26 eV (Dre91) to ≈ 4 eV (Dre86).

As another indicator of the rate of plume expansion, Figure 4.19 shows the expansion of the peak aluminum neutral line density (after it has separated from the target surface) as a function of time. The points on the graph all fall within the average velocity range of 0.24 cm/ μ s -

0.54 cm/ μ s, which is only somewhat slower than the expansion of the $1x10^{14}$ cm⁻² equicontour line.

It is important to clarify again that the perpendicular expansion velocities (v_z) are larger than the parallel expansion velocities (v_x, v_y) , as seen in the holograms, and that even the two parallel expansion velocities (v_x, v_y) are different from each other, as seen by the shape of the deposited thin film. Since the dye laser passes parallel to the target surface, it is the ensemble, parallel expansion velocity distribution that is desired for the fringe shift interpretation calculations.

Dreyfus measured the perpendicular expansion temperature (corresponding to v_z) to be in the range 0.26 eV - 4.0 eV (Dre86, Dre91). The perpendicular expansion temperature range measured here, 2.1 eV - 3.7 eV, is consistent with Dreyfus' measurements, and the parallel expansion temperature for the ablation plumes in this work is presumed to be in the range 0.1 eV - 1.0 eV. Since the plume temperature was not measured directly for the ablation plumes in this work, Dreyfus' lower value, 0.3 eV, was chosen as the value to use in the fringe shift interpretation calculations. Finally, it was shown in Chapter 2 that the resonant hologram interpretation calculation is insensitive to temperature for temperatures within the range 0.1 eV - 1.0 eV.

Theoretically Modeling Plume Expansion In Vacuum And Low Pressure, Inert, Background Gas

Laser ablation plume expansion in vacuum (or collisionless transport of the plume through the background gas) has been modeled

theoretically by Utterback, et al. (Utt76), Talents (Tal83), and Kools, et al. (Koo92) to have the form

$$n(x, y, z) = \frac{N_{o} \left(\frac{m}{2\pi k_{B}}\right)^{3/2}}{\left(T_{x}T_{y}T_{z}\right)^{1/2} (t^{3})} Exp\left\{-\frac{m}{2k_{B}t^{2}}\left[\frac{x^{2}}{T_{x}} + \frac{y^{2}}{T_{y}} + \frac{(z - ut)^{2}}{T_{z}}\right]\right\} (4.4)$$

where n is the particle density, t is time, T_x , T_y , and T_z are the expansion temperatures in the center of mass frame of reference, and u is the center of mass velocity. Integration over the x-direction (the direction of travel for the dye laser beam) gives

$$\overline{nl}(y,z) = \frac{N_{o}\left(\frac{m}{2\pi k_{B}}\right)}{\left(T_{y}T_{z}\right)^{\frac{1}{2}}(t^{2})} Exp\left\{-\frac{m}{2k_{B}t^{2}}\left[\frac{y^{2}}{T_{y}} + \frac{(z-ut)^{2}}{T_{z}}\right]\right\}$$
(4.5)

where \overline{n} is the line-integrated-particle-density. The plume expansion in Figures 4.6, 4.8, 4.12, 4.14 and 4.16 can be fit to this equation by assuming y=0 along the line of symmetry of the plume, rearranging the terms and taking the natural logarithm to obtain the form

$$\ln\left[\overline{nl} t^{2}\right] = A\left(\frac{z}{t}\right)^{2} + B\left(\frac{z}{t}\right) + C$$
(4.6)

where z is the maximum expansion distance of the equicontour with a line density \overline{nl} ,

$$A = \frac{-m}{2k_BT_z} \qquad B = -2 Au \quad and \quad C = \ln\left[\frac{m N_o}{2\pi k_B\sqrt{T_yT_z}}\right] + Au^2 \quad (4.7)$$

When $\ln[\overline{nl}*t^2]$ vs (z/t) is graphed for a given hologram, the points are binomially fit for the parameters A, B and C, from which the parameters T_z, u, and $N \sqrt{\sqrt{T_y}}$ can be calculated. While the graphs of $\ln[\overline{nl}*t^2]$ vs (z/t) have been relegated to Appendix 6, the latter three parameters are graphed versus time in Figures 4.20, 4.21 and 4.22.

First of all, it is necessary to have at least three equicontour lines for each hologram in order to fit a binomial curve to its graph of $\ln[\overline{nl}*t^2]$ vs (z/t). Three of the five time scans only had one hologram that met this requirement and so these time scans only have one point each in Figures 4.20, 4.21 and 4.22.

Secondly, it is important to remember that the parameters presented in these figures are not directly measured, they are only inferred from the analysis described above and could be used to recreate a theoretical surface plot with a similar appearance to the presented hologram tracings. Third, the expansion model assumes a point source, while the source of ablated particles in this experiment is about $0.05 - 0.10 \text{ cm}^2$.

Fourth, notice that the inferred center of mass velocity decreases quickly in time and is even negative for two holograms, while the inferred plume expansion temperature, T_z , increases rapidly in time. These results conflict with the previously made assumptions that the plume is collisionless with the background gas (and thus the center of



Fig. 4.20 Inferred center of mass velocity, u, versus time for vacuum and lower pressure argon gas backgrounds.



Fig. 4.21 Inferred, perpendicular plume expansion temperature, T_z , versus time in vacuum and lower pressure argon gas backgrounds.



Fig. 4.22 The inferred parameter, $N_0 x T_y^{-1/2}$, versus time in vacuum and lower pressure argon gas backgrounds.



Fig. 4.23 Inferred plume expansion temperature, T_y , versus time in vacuum and lower pressure argon gas backgrounds. (Combination of Fig. 4.18(a) and Fig. 4.22.)

mass velocity does not decrease rapidly) and that the plume temperature (velocity distribution) is not changing with time because, again, the plume is collisionless during the time of investigation. Instead, these graphs are an indication that the plume expansion model does not fit the holographic data well. In practice, there probably exist two groups of particles (components) in the ablation plume and the above model only describes the motion of one of those components.

One group of particles (ions and charge exchange neutrals accelerated by the ambipolar potential) energetically moves away from the target and undergoes self-similar expansion according to Equation 4.5 while the other group of particles stays very close (within a few mm) to the target surface. Hereafter, these will be called the expanding plume and the residual plume for simplicity.

The expanding plume and the residual plume have been observed and discussed for laser ablation in vacuum by Geohegan using emission spectroscopy (Geo92, Geo93) and by Bakos, et al. using TOFPRL (Bak87). The phenomena is also visible in the PLIF data discussed by Cappelli, et al. (Cap90). However, this should not be confused with discussions distinguishing between species within the expanding plume such as Saenger, who discussed the differential expansion of the four different kinds of atoms in the ablation of YBaCuO (Sae91), or Kools, et al., who observed a CuCl expanding plume and a slower, Cu₃Cl₃ expanding plume following the ablation of copper in the presence of chlorine gas (Koo92).

Therefore, the holograms in Figures 4.6, 4.8, 4.12, 4.14 and 4.16 represent the RHI-detectable portion of the sum of the expanding plume and the residual plume. In most cases, however, the expanding plume is

not detectable after separation from the residual plume. For example, Fig. 4.12(d) clearly shows a dominant, expanding plume at 1.00 μ s, while at 2.00 μ s in Fig. 4.12(f), the expanding plume is no longer distinct and the most dense region is next to the target. The single exception to this in the data presented thus far is Fig. 4.14(h), where the residual plume is no longer detectable and a detectable expanding plume has separated from the target at 1.20 μ s.

For completeness, the parameter $N_o/\sqrt{T_y}$ is graphed versus time in Fig. 4.22 and Ty is then calculated using the values of N₀ presented in Fig. 4.18(a) and graphed versus time in Fig. 4.23. As presented above, caution should be used when interpreting these numbers, especially because $N_o/\sqrt{T_y}$ is proportional to Exp(C) as given in Eqn. 4.7.

Laser Ablation In Higher Pressure, Inert, Background Gas

Figure 4.24 shows a RHI time scan in 1 Torr argon gas. The first thing to notice is that with a high enough background pressure, a shock wave was observed at 1.02 μ s and 4.00 μ s. Previously, DLRAP studies have only detected shock waves down to pressures of ≈ 5 Torr (Ven92). It is presumed that a shock wave is present at the other times, but that for some reason it was not detected as clearly as in Figs. 4.24(a) and (f). Secondly, notice that the plume expansion has slowed considerably with respect to the previous time scans at lower pressures.

Third, note the mushroom shape of the ablation plume. Sappey and Gamble attribute this shape for copper ablation plumes to the exothermic



Fig. 4.24 Time scan in 1 Torr argon gas. Hologram (a) taken at $(1.02 \ \mu s, -0.006 \ nm, \approx 2.3 \ J/cm^2$) and hologram (b) taken at $(2.02 \ \mu s, -0.006 \ nm, \approx 2.4 \ J/cm^2$). (c) is an equicontour plot of (a), dashed line corresponds to a shock wave and solid line corresponds to (0.5 FS =7.0x10¹³ cm⁻²); (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS =7.0x10¹³ cm⁻²), (1.0 FS=1.4x10¹⁴ cm⁻²), (1.5 FS =2.1x10¹⁴ cm⁻²), (2.0 FS =2.9x10¹⁴ cm⁻²), (2.5 FS=3.6x10¹⁴ cm⁻²), and (3.0 FS=4.4x10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.24 Time scan in 1 Torr argon gas, continued. Hologram (e) taken at (2.50 μ s, -0.006 nm, $\approx 2.1 \text{ J/cm}^2$) and hologram (f) taken at (4.00 μ s, -0.006 nm, $\approx 2.4 \text{ J/cm}^2$). (g) is an equicontour plot of (e); lines correspond to increasing fringe shifts, from outer to inner line: (1.0 FS =1.4x10¹⁴ cm⁻²), (2.0 FS=2.9x10¹⁴ cm⁻²), (3.0 FS=4.4x10¹⁴ cm⁻²), and (4.0 FS=6.0x10¹⁴ cm⁻²). (h) is an equicontour plot of (f); dashed line corresponds to a shock wave and solid lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS =7.0x10¹³ cm⁻²), (1.0 FS =1.4x10¹⁴ cm⁻²), (1.5 FS =2.1x10¹⁴ cm⁻²), (2.0 FS =2.9x10¹⁴ cm⁻²), (2.5 FS=3.6x10¹⁴ cm⁻²), (3.0 FS=4.4x10¹⁴ cm⁻²), (3.5 FS=5.2x10¹⁴ cm⁻²) and (4.0 FS=6.0x10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.24 Time scan in 1 Torr argon gas, continued. Hologram (i) taken at (10.0 μ s, -0.006 nm, ~2.4 J/cm²) and hologram (j) taken at (20.0 μ s, -0.006 nm, ~2.4 J/cm²). (k) is an equicontour plot of (i), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS =7.0x10¹³ cm⁻²), (1.0 FS =1.4x10¹⁴ cm⁻²), (1.5 FS =2.1x10¹⁴ cm⁻²), and (2.0 FS =2.9x10¹⁴ cm⁻²). (l) is an equicontour plot of (j); line corresponds to (0.5 FS =7.0x10¹³ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.24 Time scan in 1 Torr argon gas, continued. Hologram (m) taken at $(50.0 \,\mu\text{s}, -0.006 \text{nm}, \approx 2.4 \,\text{J/cm}^2)$. (n) is an equicontour plot of (m), line corresponds to $(0.5 \,\text{FS} = 7.0 \times 10^{13} \,\text{cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)

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Fig. 4.25(a) Total number of aluminum neutral atoms, N $_{\rm o}$, versus time in 1 Torr argon gas.



Fig. 4.25(b) Expansion of the 1×10^{14} cm⁻² equicontour line in 1 Torr argon gas.

condensation of the atoms to dimers and to the fluid dynamics of one directionally expanding fluid in another (Sap92). Fourth, notice that the plume separates from the target by 20 μ s. Finally, notice that the plume is observed as late as 50.0 μ s with RHI at 1 Torr, while the latest time an ablation plume was observed in the previous vacuum data was 4.0 μ s in a DLRAP photo, which is more sensitive than RHI. Not only does this indicate collisions with the argon gas are holding back the expanding ablation plume, but also that the conditions are such in the plume to prevent significant condensation of the aluminum neutral atoms after that amount of time (50 μ s).

As with the holographic time scans at lower pressures, the peak line-density in the earliest holograms is obscured by absorption. The peak line-density at 2.50 µs and 4.00 µs is 6.0×10^{14} cm⁻², and it decreases to the detectability limit at this $|\Delta\lambda|$, 7×10^{13} cm⁻², by 20.0 µs. These correspond to approximate aluminujm neutral densities of $\approx 7 \times 10^{14}$ cm⁻³ and 7×10^{13} cm⁻³ for the earlier and later times, respectively.

Figure 4.25(a) shows the total number of aluminum neutral atoms counted for each hologram versus time. The value of N₀, $\approx 3x10^{13}$ to $\approx 3x10^{14}$, is similar to the values counted for the other time scans and the value clearly decreases with time as the aluminum atoms expand outward beyond the range of detection using RHI. Figure 4.25(b) shows the expansion of the $1x10^{14}$ cm⁻² equicontour as a function of time. The expansion velocity of 0.36 cm/µs is nearly half the slowest expansion velocity for the vacuum and lower pressure time scans (0.77 cm/µs). The expansion of the $1x10^{14}$ cm⁻² equicontour then appears to more or less stabilize at about 1 cm between 5 µs and 50 µs. Figure 4.26 shows a RHI time scan in 1 Torr argon RF-plasma. The RF-plasma was produced between two parallel plates enclosing the region of plume expansion, as described in Chapter 3. The argon plasma is assumed to have a maximum electron density of 1×10^{10} cm⁻³ and a maximum kinetic temperature of 2 eV (Bro93). The Stark broadening introduced into the aluminum neutral line-density calculations due to the RF-plasma is minimal for the given electron density and temperature. Therefore, the calculations in Chapter 2 assuming no Stark broadening are used here for the plasma case as well.

The first thing to notice is that the plume shapes look very similar to the non-plasma case. Secondly, note the separation of the expanding plume and the residual plume at 20.0 μ s in Fig. 4.26(1) and only the expanding plume at 50.0 μ s in Fig. 4.26(n). Third, the counted number of aluminum atoms in the plume is slightly less for the RF-plasma case than for the non-plasma case, as shown in Fig. 4.27(a). Fourth, the 1x10¹⁴ cm⁻² equicontour line appears to expand farther away from the target in the RF-plasma case than in the gas case.

At this time, it is unknown whether the observed differences are: 1) just shot-to-shot, statistical variations of the ablation plumes; 2) caused by the RF electric field directly affecting the ablated aluminum ions at very early times (< 0.1 μ s); or 3) caused by collisions between the ablation plume and the energetic electrons and argon ions during the entire expansion time. For example, energetic collisions of the argon plasma with the ablation plume could cause either ionization or condensation of the aluminum atoms, thus causing the observed reduction in the number of neutral atoms. Since this issue is the focal point of Chi Hong Ching's dissertation, which is forthcoming, Fig. 4.26 is shown here as an example



Fig. 4.26 Time scan in 1 Torr argon RF/plasma. Hologram (a) taken at (1.00 μ s, -0.006 nm, $\approx 2.5 \text{ J/cm}^2$) and hologram (b) taken at (2.00 μ s, -0.006 nm, $\approx 2.3 \text{ J/cm}^2$). (c) is an equicontour plot of (a), dashed line corresponds to a shock wave and solid line corresponds to (0.5 FS =7.0x10¹³ cm⁻²); (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: (1.0 FS=1.4x10¹⁴ cm⁻²), (2.0 FS =2.9x10¹⁴ cm⁻²), (3.0 FS=4.4x10¹⁴ cm⁻²), and (4.0 FS=6.0x10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.26 Time scan in 1 Torr argon RF/plasma, continued. Hologram (e) taken at $(2.50 \,\mu\text{s}, -0.006 \,\text{nm}, \approx 2.1 \,\text{J/cm}^2)$ and hologram (f) taken at $(4.02 \,\mu\text{s}, -0.006 \,\text{nm}, \approx 2.4 \,\text{J/cm}^2)$. (g) is an equicontour plot of (e), and (h) is an equicontour plot of (f); lines correspond to increasing fringe shifts, from outer to inner line: $(0.5 \,\text{FS} = 7.0 \times 10^{13} \,\text{cm}^{-2})$, $(1.0 \,\text{FS} = 1.4 \times 10^{14} \,\text{cm}^{-2})$, $(1.5 \,\text{FS} = 2.1 \times 10^{14} \,\text{cm}^{-2})$, $(2.0 \,\text{FS} = 2.9 \times 10^{14} \,\text{cm}^{-2})$, $(2.5 \,\text{FS} = 3.6 \times 10^{14} \,\text{cm}^{-2})$ and $(3.0 \,\text{FS} = 4.4 \times 10^{14} \,\text{cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)

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Fig. 4.26 Time scan in 1 Torr argon RF/plasma, continued. Hologram (i) taken at (10.0 μ s, -0.006 nm, =2.4 J/cm²) and hologram (j) taken at (20.0 μ s, -0.006 nm, =2.4 J/cm²). (k) is an equicontour plot of (i), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS =7.0x10¹³ cm⁻²), (1.0 FS =1.4x10¹⁴ cm⁻²), and (1.5 FS =2.1x10¹⁴ cm⁻²). (l) is an equicontour plot of (j); line corresponds to (0.5 FS =7.0x10¹³ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.26 Time scan in 1 Torr argon RF/plasma, continued. Hologram (m) taken at $(50.0 \,\mu\text{s}, -0.006 \text{nm}, \approx 2.4 \,\text{J/cm}^2)$. (n) is an equicontour plot of (m), line corresponds to $(0.5 \,\text{FS} = 7.0 \times 10^{13} \text{cm}^{-2})$. (From laser spot #2 of Fig. 3.2(b).)

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Fig. 4.27(a) Total number of aluminum neutral atoms, N_o , versus time in 1 Torr argon RF/plasma.



Fig. 4.27(b) Expansion of the 1×10^{14} cm⁻² equicontour line in 1 Torr argon RF/plasma.

of the copious laser-ablation-plume-expansion-in-plasma-environments data by which this author has demonstrated the applicability of RHI.

Figure 4.28 shows an early time scan at a background argon gas pressure of 35 Torr. Although a complete time scan at this pressure is not presented, the holograms show two interesting effects. First of all, this is the only pressure at which non-resonant fringe shift was observed and secondly, the plume expansion is even slower than at 1 Torr. Although the Lorentz broadening mechanisms become on the order of the natural broadening for a background pressure of 35 Torr, the effect of the overall broadening induced onto the calculation of line-densities is minimal. Therefore, the calculations of aluminum neutral line-density interpretation in Chapter 2, which assume no Lorentz broadening, are utilized in the interpretation of the 35 Torr data.

The only nonresonant fringe shift was observed at very early times, 0.2 μ s and 0.5 μ s. Figures 4.28(a) and (b) show a near-resonant and a non-resonant hologram, respectively, taken at 0.22 μ s. Similarly, Figures 4.28(e) and (f) show a near-resonant and a non-resonant hologram, respectively, taken at 0.52 μ s. The resonant-only fringe shift for each time was found by overlaying the two traces for the same time and subtracting the fringe shifts at each spatial location. Figures 4.28(i) and (j) are therefore the resulting subtractions at 0.22 μ s and 0.52 μ s, respectively, and represent the resonant-only fringe shift due to interaction with aluminum neutral atoms. Figures 4.28(k) to (n) show holograms and their tracings at 1.80 μ s and 3.84 μ s. There was no fringe shift observed at the off-resonant wavelength, so all of the observed fringe shift is due to resonant effects.

The total number of aluminum neutral atoms counted at 0.22 μ s is



Fig. 4.28 Time scan in 35 Torr argon gas. Hologram (a), taken at $(0.22 \ \mu\text{s}, +0.004 \text{nm}, \approx 2.4 \ \text{J/cm}^2)$, is the near-resonance hologram for 0.22 $\ \mu\text{s}$. Hologram (b), taken at $(0.22 \ \mu\text{s}, +0.503 \ \text{nm}, \approx 2.4 \ \text{J/cm}^2)$, is the off-resonance hologram for 0.22 $\ \mu\text{s}$. (c) is an equicontour plot of (a), lines correspond to increasing fringe shifts, from outer to inner line: (1.0 FS), (2.0 FS), (3.0 FS), (4.0 FS), and (5.0 FS); (d) is an equicontour plot of (b), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS), (1.0 FS), (1.5 FS), (2.0 FS). The subtraction of (d) from (c) is shown in (i). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.28 Time scan in 35 Torr argon gas, continued. Hologram (e), taken at $(0.52 \ \mu\text{s}, +0.004 \text{nm}, \approx 2.4 \ \text{J/cm}^2)$, is the near-resonance hologram for 0.52 $\ \mu\text{s}$. Hologram (f), taken at $(0.52 \ \mu\text{s}, +0.503 \ \text{nm}, \approx 2.4 \ \text{J/cm}^2)$, is the off-resonance hologram for 0.52 $\ \mu\text{s}$. (g) is an equicontour plot of (e), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS), (1.0 FS) and (1.5 FS); (h) is an equicontour plot of (f), line corresponds to (0.5 FS). The subtraction of (h) from (g) is shown in (j). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.28 Time scan in 35 Torr argon gas, continued. Tracing (i) is the point-to-point subtraction of (d) from (c) and represents the resonant-only fringe shift at 0.22 μ s. Lines correspond to increasing fringe shifts, from outer to inner line: (1.0 FS=9.0×10¹³ cm⁻²), (2.0 FS=1.9×10¹⁴ cm⁻²), and (3.0 FS =3.0×10¹⁴ cm⁻²). Tracing (j) is the point-to-point subtraction of (h) from (g) and represents the resonant-only fringe shift at 0.52 μ s. Lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS=4.3×10¹³ cm⁻²) and (1.0 FS=9.0×10¹³ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



Fig. 4.28 Time scan in 35 Torr argon gas, continued. Hologram (k) taken at (1.80 μ s, +0.004nm, ≈2.3 J/cm²) and hologram (l) taken at (3.84 μ s, +0.004 nm, ≈2.5 J/cm²). (m) is an equicontour plot of (k), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS =4.3x10¹³ cm⁻²), (1.0 FS=9.0x10¹³ cm⁻²) and (1.5 FS=1.4x10¹⁴ cm⁻²). (n) is an equicontour plot of (l), lines correspond to increasing fringe shifts, from outer to inner line: (0.5 FS=4.3x10¹³ cm⁻²), (1.0 FS=9.0x10¹³ cm⁻²), (1.0 FS=9.0x10¹³ cm⁻²), (1.0 FS=1.4x10¹⁴ cm⁻²), and (2.0 FS=1.9x10¹⁴ cm⁻²). (From laser spot #2 of Fig. 3.2(b).)



2.5x10¹² cm⁻². The value of N₀ is an order of magnitude smaller than the values of N₀ determined for previous time scans. This is primarily due to the inadequate spatial resolution of the holograms at early times when the plume is very close to the target surface and could also be due to forced recondensation on the target at these pressures. It is suggested that a hologram at a much later time, such as 20 μ s, would provide a much better count of the total number of ablated atoms in the plume.

Figure 4.29 shows the expansion of the 1×10^{14} cm⁻² equicontour. The initial expansion velocity of ≈ 0.084 cm/µs is about 4 times smaller than the expansion velocities measured at 1 Torr and the expansion velocity decreases after 0.5 µs to about 0.01 cm/µs.

Fig. 4.30 shows the total number of aluminum neutral atoms in the ground state measured for all eight time scans; the values range from 2.5×10^{12} to 3.2×10^{14} . Since many of the earlier N₀'s in each time scan



Fig. 4.30(a) Total number of aluminum neutral atoms, N_0 , versus time in vacuum and argon background gas; early times.



Fig. 4.30(b) Total number of aluminum neutral atoms, N_o , versus time in vacuum and argon background gas; later times.

are >1x10¹⁴ and since $3x10^{14}$ was the estimated maximum number of particles in the laser ablation deposition discussion above, $N_0 \approx 3x10^{14}$ is a good estimate of the number of aluminum ground-state neutral atoms that the KrF laser is ablating from the aluminum targets for the experimental parameters in this work.

Resonant refraction is very strong, allowing only the edges of the ablation plume to be analyzed. The peak line-density for a given hologram was $3x10^{14}$ cm⁻² at 0.22 µs and $2x10^{14}$ cm⁻² at 3.84 µs. Considering the small size of the plume, these correspond to particle densities of $1x10^{15}$ cm⁻³ and $6x10^{14}$ cm⁻³, respectively.

Figure 4.31 shows the expansion of the 1×10^{14} cm⁻² equicontour line for aluminum neutrals in 1 Torr argon gas and plasma, 35 Torr argon gas, and the deposition-in-vacuum plume expansion case for comparison. Notice that the plume expansion is significantly slowed for the 1 Torr case and even more so for the 35 Torr case. The approximate expansion velocities for all of the time scans are compiled in Figure 4.32 versus pressure. For 210 mTorr and less, the expansion velocity is in the range 1.0 cm/ μ s - 1.4 cm/ μ s, while for higher background pressures of argon the expansion velocity decreases to as low as 0.084 for the 35 Torr case.



Fig. 4.31(a) Expansion of the 1×10^{14} cm⁻² equicontour line in 1 Torr and 35 Torr argon; early times. Deposition in vacuum expansion shown for comparison (from fig. 4.18(b)).



Fig. 4.31(b) Expansion of the 1×10^{14} cm⁻² equicontour line in 1 Torr and 35 Torr argon; later times. Deposition in vacuum expansion shown for comparison (from fig. 4.18(b)).



CHAPTER 5 CONCLUSIONS

Resonant Holographic Interferometry (RHI) has been demonstrated to measure the expanding absolute-line-density profiles of aluminum neutral atoms caused by KrF excimer laser ablation of aluminum targets using a dye laser tuned near the Al(I), 394.401 nm, resonant transition. Dye-Laser-Resonance-Absorption-Photography (DLRAP), which is more sensitive than RHI, has also been utilized to observe the ablation plumes under similar conditions. Plume expansion was studied for KrF excimer laser ablation (248 nm) of aluminum ($\approx 0.1 \text{ cm}^2$, 2 - 6 J/cm²) in vacuum and in background argon gas pressures of 14 mTorr, 52 mTorr, 210 mTorr, 1 Torr, and 35 Torr.

The theory developed by Measures (Mea70) to calculate absoluteline-densities from resonant interferometry, which accounted for Doppler shift effects, has been extended to also account for the diagnostic laser bandwidth and the selective absorption of the diagnostic laser by the resonant medium. The results show that for lower line-densities where there is little absorption, Doppler shift effects dominate, causing the observed fringe shift to increase with respect to the value determined from the simplest interferogram interpretation model (Gre64) and as Measures had predicted. For higher line-densities however, selective absorption effects dominate, causing the observed fringe shift to decrease with respect to the simplest interferogram interpretation model. The results also show that fringe shift interpretation is not very sensative to the parallel expansion temperature for temperatures in the range 0.1 eV - 1.0 eV.

Absolute-line-densities for the aluminum neutral atoms have been measured in the range 4.3×10^{13} cm⁻² to 1.0×10^{15} cm⁻², which imply aluminum neutral particle densities up to 1×10^{15} cm⁻³. The total number of aluminum neutral atoms in a plume has been measured to be $\approx 3 \times 10^{14}$. This value corresponds to a surface etch rate of 1 nm/pulse, which is close to the etching rate of Kelly and Rothenburg for 2.3 J/cm² ablation of aluminum (Kel85). Knowing the reflectivity of aluminum to be 92.5%, the energy absorbed by the target per ablated particle was estimated to be ≈ 466 eV, where most of the absorbed energy is presumably dissipated by the aluminum target. This is consistent with the theory that the primary ablation mechanism for metals is thermal ablation (Dre91), where the absorbed laser energy raises the surface temperature causing liquefication and vaporization of the target surface.

The observed ablation plume expansion in vacuum and for pressures ≤ 210 mTorr was consistent with the two-component theory for laser ablation: a fast component of ablated particles, corresponding to ions and charge exchange neutrals accelerated by the ambipolar potential, moved quickly away from the target surface while a slow component of particles lingered near the surface. This expansion model has been discussed and observed by Geohegan using emission photography (Geo92, Geo93) and by Bakos, et al. using TOFPRL (Bak87). Two component expansion is also visible in the PLIF data presented by Cappelli, et al. (Cap90). For most of the holograms in this work, the fast component was at or below the edge of RHI-detectability at the time of separation with the slow component.

Expansion velocities for the 1×10^{14} cm⁻² equicontour line for pressures ≤ 210 mTorr were in the range 1.05 cm/µs to 1.41 cm/µs. This is consistent with other recent measurements of the TOF most-probable expansion velocity (Geo89, Wan91, Zhe89) and the leading edge expansion velocity (Dre91, Ven92a, Wal86) of laser ablation expansion in vacuum. Expansion velocities in the range of 0.24 cm/µs - 0.54 cm/µs were measured for the center of the plume. The perpendicular expansion velocities measured here correspond to kinetic temperatures in the range 2 eV to 4 eV, which is within the range measured by Dreyfus (Dre86, Dre91) for the ablation of copper under similar conditions.

The expansion of the ablation plume in 1 Torr and 35 Torr argon gas backgrounds was observed to be slower than at pressures less than 210 mTorr; ≈ 0.34 cm/µs was measured for 1 Torr and ≈ 0.084 cm/µs was measured for 35 Torr. Hydrodynamic effects were also observed, such as mushroom cloud shaped plumes. Sappey and Gamble attribute the mushroom shape for their copper ablation plumes to the exothermic condensation of the atoms to dimers and to the fluid dynamics of one directionally expanding fluid in another (Sap92).

Plume expansion into a 1 Torr argon, RF-plasma was compared with the expansion into a 1 Torr argon gas. It was observed that the ablation plume appeared to expand faster in the plasma and that the total number of aluminum neutrals was lower in the plasma. At this time, it is unknown whether the observed differences are: 1) just statistical variations of the ablation plumes; 2) caused by the RF electric field directly affecting the ablated aluminum ions and electrons at very early times (< 0.1 μ s); 3) caused by collisions between the ablation plume and the energetic electrons and argon ions during the entire expansion time; or 4) some combination of the above.

Laser ablation deposition of aluminum onto amorphous silicon at room temperature was performed with a target-substrate separation of 1.5 cm. The thin film was observed to have an oval shape, with the larger dimension corresponding to the narrow dimension of the ablation laser spot. This effect has been theoretically modeled and observed in atmospheric air by Miotello, et al. (Mio92), and is confirmation that the expansion of the ablated particles in this experiment is primarily governed by self-collision and self-diffusion mechanisms, especially for expansion into vacuum and lower pressure background gases. APPENDICES

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APPENDIX 1 MEASURES' THEORY

The following is a full derivation of the theory derived from Measures (Mea70). Others who have used or built on this theory include Odenthal, et al. (Ode80), Jellison, et al.(Jel81), Kügler, et al. (Küg83), Bone, et al. (Bon86), Craig, et al. (Cra86), and Woisetschläger, et al. (Woi91, Woi92).

The classical model (Mar67) of the interaction of a photon with an atom (or molecule) is described by a sinusoidal electric field, $\underline{E} = E_0 e^{i\omega t}$, driving an elastically bound electron whose equation of motion, under a dissipative force proportional to the velocity, is

$$m\ddot{z} = -m\gamma\dot{z} - m\omega_0^2 z + eE_0 e^{i\omega t} \qquad (A1.1)$$

The solution of this equation has the form

$$z = \frac{e'_{m}}{\omega_{0}^{2} - \omega^{2} + i\gamma\omega} E = \frac{\tilde{\alpha}E}{e} , \qquad (A1.2)$$

where ω_0 is the resonance frequency of the binding force on the electron, γ is the dissipative damping constant, and $\tilde{\alpha}$ is the complex polarizability of the electrically bound electron. The quantum mechanical version of $\tilde{\alpha}$ for an atom in the ith quantum state is (Mar67)

$$\tilde{\alpha}_{i} = \frac{\begin{pmatrix} e^{2}/m \end{pmatrix} f_{ik}}{\omega_{ik}^{2} - \omega^{2} + i\gamma_{ki}\omega}, \qquad (A1.3)$$

where ω_{ik} is the resonance frequency, f_{ik} is the oscillator strength for absorption, and γ_{ki} is the damping constant (to be discussed later). The
oscillator strength of a transition is related to the Einstein A coefficient for that transition by the relationship (Bershader)

$$f_{ik} = \frac{4\pi^2 \varepsilon_0 mc^3}{\omega_{ik}^3 e^2 \hbar} A_{ik}$$
(A1.4)

For a photon passing through a collection of atoms, the collection of particles can be treated as a dielectric material. The effective electric field on the atom can then be determined using classical electromagnetic theory (Nay85, Ode80).

Figure A.1 shows the cross section of a dielectric material with dipole moment density, \vec{P} , in an external electric field, \vec{E}_{ext} , both in the \hat{z} direction. Internally, there is a cavity with atom A as well as N other identical atoms. The rest of the material is treated as a continuum. The external field induces a surface charge on the dielectric material, σ_p , and



Fig. A.1 Determining the electrical effects on an individual atom of a dielectric material polarized by an external electric field. (Nay85)

a surface charge around the cavity, σ_p' . The electric field as seen by atom A is

$$\vec{E}_a = \vec{E} + \vec{E}_{cavity} + \vec{E}_i , \qquad (A1.5)$$

where

 \vec{E} = (external field) - (dielectric surface charge field) = $\vec{E}_{ext} - \frac{\vec{P}}{\epsilon_0}$ (A1.6)

$$\vec{E}_{cavity} = + \frac{\vec{P}}{3\varepsilon_0} , \qquad (A1.7)$$

and \vec{E}_i is the field due to all the atoms in the dielectric, which is zero since the dielectric is a fluid for our case. Therefore,

$$\vec{E}_{a} = \vec{E} + \frac{\vec{P}}{3\varepsilon_{o}}$$
(A1.8)

which can also be written

$$\vec{P} = 3\varepsilon_{o}(\vec{E}_{a} - \vec{E})$$
 (A1.9)

For linear materials such as fluids, the dipole moment of the material can be written as $\vec{p} = \alpha \vec{E}_a$. For N such atoms per unit volume, one obtains

$$\vec{P} = N\vec{p} = N\alpha \vec{E}_a = N\alpha \left(\vec{E} + \frac{\vec{P}}{3\varepsilon_o}\right)$$
 (A1.10)

If the dielectric is simple, then $\vec{P} = \chi \epsilon \vec{E}$ where the electric displacement, \vec{D} , is

$$\vec{D} = \varepsilon_0 (1 + \chi) \vec{E} = \varepsilon_0 K \vec{E} = \varepsilon \vec{E}$$
(A1.11)

which gives us

$$\vec{\mathbf{P}} = \mathbf{N}\alpha \left(\vec{\mathbf{E}} + \frac{\vec{\mathbf{P}}}{3\varepsilon_0}\right) = \mathbf{N}\alpha \left(\frac{\vec{\mathbf{P}}}{\chi\varepsilon_0} + \frac{\vec{\mathbf{P}}}{3\varepsilon_0}\right)$$
 (A1.12)

and finally leads to the form

$$\frac{N\alpha}{3\varepsilon_{o}} = \frac{\chi}{\chi + 3} = \frac{K - 1}{K + 2}$$
(A1.13)

The index of refraction is defined as

$$n \equiv \frac{c}{v} = \sqrt{KK_m}$$
(A1.14)

In most materials, $K_m \cong 1$ and so $n^2 \cong K$. Using the complex index of refraction, where $|\tilde{n}| \cong n \equiv 1$ so that $\tilde{n}^2 \cong K$, equation A1.13 becomes

$$\frac{N\alpha}{3\varepsilon_o} \cong \frac{\tilde{n}^2 - 1}{\tilde{n}^2 + 2} \quad . \tag{A1.15}$$

Next, assuming that the dielectric gas in question consists of different species having different polarizabilities $\tilde{\alpha}_i$ and densities N_i , equation A1.15 becomes

$$\frac{\tilde{n}^2 - 1}{\tilde{n}^2 + 2} = \frac{1}{3\varepsilon_0} \sum_{i} \tilde{\alpha}_i N_i = \frac{1}{3\varepsilon_0} \sum_{i} \frac{N_i f_{ik} (e^2/m)}{\omega_{ik}^2 - \omega^2 + i\gamma_{ki} \omega}$$
(A1.16)

using A1.3.

In truth, this is a simplification of the quantum mechanical events occurring in the atom during an electronic transition. As an electron changes electronic states, the whole electron cloud surrounding the atom changes shape. In turn, the induced polarizability of the atom and therefore the index of refraction change as well. Bershader gives the correct form as (Ber71)

$$\frac{\tilde{n}_{j}^{2}-1}{\tilde{n}_{j}^{2}+2} = \frac{N_{j}e^{2}}{3m\varepsilon_{o}} \left\{ \sum_{k>j} \frac{f_{jk}}{\omega_{jk}^{2}-\omega^{2}+i\gamma_{kj}\omega} - \sum_{i(A1.17)$$

For the derivation here, the second term can be ignored for two reasons. First, this work investigates the index of refraction very near a single, relatively isolated transition. Second, for the Al(I) $3^2P_{1/2} - 4^2S_{1/2}$,

394.4 nm, transition under investigated, there are no electronic transitions to states lower than $3^{2}P_{1/2}$.

So, now assuming one transition of interest, equation A1.16 can be simplified by assuming $\tilde{n}^2 - 1 \approx 2(\tilde{n} - 1)$, $\tilde{n}^2 + 2 \approx 3$, $\omega_0 \equiv \omega_{ik}$, $\omega_0 \approx \omega$, $r_0 = e^2 / (4\pi\epsilon_0 mc^2) = 2.818 \times 10^{-13} cm$ and

$$\omega_{o}^{2} - \omega^{2} + i\gamma\omega = (\omega_{o} + \omega)(\omega_{o} - \omega) + i\gamma\omega \equiv \omega_{o}[2(\omega_{o} - \omega) + i\gamma] \quad (A1.18)$$

and thus giving

$$\tilde{n} - 1 = \frac{N_i f_{ik} \pi r_o c^2}{\omega_o} \frac{\left\{ (\omega_o - \omega) - i \left(\frac{\gamma}{2}\right) \right\}}{(\omega_o - \omega)^2 + \frac{\gamma^2}{4}}$$
(A1.19)

Following Marlow (Mar67) and Measures (Mea70), the damping constant γ is defined as the sum of the natural, Lorentz, Holtzmark and Stark line-broadenings. The natural broadening term is defined as (Bra92, Ber71)

$$\gamma_{\rm N} = \sum_{\rm m} [A_{\rm im} + A_{\rm km}] \qquad (A1.20)$$

where the summation is over all of the states m to which an electron can relax from states i and k for the transition in question, i-k. The sum of the Lorentz and Holtzmark broadenings is given as (Bra92)

$$\gamma_{\rm L} + \gamma_{\rm H} = \sum_{\rm m} 3.37 \, {\rm x} \, 10^{10} \left[{\rm d}_{\rm i}({\rm A}) \right]^2 \sqrt{\frac{1}{{\rm M}_{\rm i(amu)}} + \frac{1}{{\rm M}_{\rm m(amu)}}} \left(\frac{{\rm P}_{\rm m}({\rm atm})}{\sqrt{{\rm T}_{\rm m}({\rm }^{\rm o}{\rm k})}} \right) ({\rm A}1.21)$$

where the summation is over all gas species m that are present, including the species of interest, i, and d_i is the effective hard-sphere diameter of the species of interest. The Stark broadening is given as (Bra92)

$$\gamma_{\rm S} = 2\pi \left(3*10^{18} \frac{\mathring{\rm A}}{\rm S}\right) 8.19*10^{-19} \left(1-0.7 N_{\rm D}^{-\frac{1}{3}}\right) \left(n_{\rm i}^2 - n_{\rm k}^2\right) \left(\frac{Z_{\rm p}^{\frac{1}{3}}}{Z_{\rm e}}\right) \left[N_{\rm e}(\rm cm^{-3})\right]^{\frac{2}{3}}$$
(A1.22)

where

$$N_{\rm D} = \frac{4\pi}{3} N_{\rm e} \lambda_{\rm D} = 1.72 * 10^9 \frac{[T(ev)]^{3/2}}{[N_{\rm e}(cm^{-3})]^{1/2}}$$
(A1.23)

 Z_p is the charge on the ion, Z_e is the atomic number, and n_i and n_k are the principle quantum numbers of the transition i-k.

Doppler broadening is accounted for assuming that the medium is a neutral gas in Maxwell-Boltzman equilibrium. The Doppler shifted frequency that each atom sees is

$$\omega' = \omega \sqrt{\frac{1+\beta}{1-\beta}} = \omega (1+\beta)^{1/2} (1-\beta)^{-1/2} \equiv \omega \left(1+\beta+\frac{1}{2}\beta^2+...\right)$$
(A1.24)

and thus

$$\omega'_{o} \cong \omega_{o} \left(1 + \frac{v_{x}}{c}\right)$$
(A1.25)

where the \hat{x} -direction is defined as the direction the diagnostic photons are traveling.

Next, replace N in equation A1.19 with the Maxwell-Boltzman integral

$$N'_{i} = N_{i} \left(\frac{m}{2\pi k_{B}T}\right)^{3/2} \int_{-\infty}^{\infty} \int dv_{x} dv_{y} dv_{z} \exp\left(-\frac{m}{2k_{B}T}\left[v_{x}^{2} + v_{y}^{2} + v_{z}^{2}\right]\right)$$

(A1.26)

and integrate over dv_y and dv_z to get the 1-D Maxwellian

$$N'_{i} = N_{i} \left(\frac{m}{2\pi k_{B}T}\right)^{1/2} \int_{-\infty}^{\infty} dv_{x} \exp\left(\frac{-mv_{x}^{2}}{2k_{B}T}\right) \qquad (A1.27)$$

Dropping the subscript x, equation A1.19 is now written as

$$\tilde{n} - 1 \cong \left(\frac{\pi r_{o}c^{2}N_{i}f_{ik}}{\omega_{o}}\right) \left(\frac{m}{2\pi k_{B}T}\right)^{1/2} \int_{-\infty}^{\infty} dv \exp\left(\frac{-mv^{2}}{2k_{B}T}\right) \left(\frac{\omega_{o}\left(1 + \frac{v}{c}\right) - \omega\right) - i\left(\frac{\gamma}{2}\right)}{\left(\omega_{o}\left(1 + \frac{v}{c}\right) - \omega\right)^{2} + \frac{\gamma^{2}}{4}}$$
(A1.28)

Measures now uses the following simplifications

$$\beta = \frac{\omega_o}{c} \sqrt{\frac{2k_BT}{m}} \qquad \alpha = \frac{\gamma_{ik}}{2\beta} \qquad u = \frac{\omega - \omega_o}{\beta} \qquad y = \frac{v\omega_o}{c\beta} \qquad (A1.29)$$

giving

$$\tilde{n} - 1 \equiv \left(\frac{\sqrt{\pi} r_{o} c^{2} N_{i} f_{ik}}{\beta \omega_{o}}\right)_{-\infty}^{\infty} dy \frac{(y - u - i\alpha) exp(-y^{2})}{(y - u)^{2} + \alpha^{2}}$$

$$\equiv \left\{\frac{c}{\omega_{o}}\right\} - i \left(\frac{\pi^{3/2} r_{o} c f_{ik}}{\beta}\right) (N_{i}) \left[\frac{i}{\pi} \int_{-\infty}^{\infty} dy \frac{exp(-y^{2})}{y - [u - i\alpha]}\right]$$

$$\equiv \left\{\frac{c}{\omega_{o}}\right\} (-iA) N_{i} W (-u + i\alpha)$$
(A1.30)

where (Abr72)

$$W(z) = \exp(-z^2) * (1 + erf{iz})$$
 (A1.31)

and where $[i\sqrt{\pi} W(z)]$ is the plasma dispersion function.

Using the convention

$$\underline{\mathbf{E}} = \mathbf{E}_{o} \exp(i\omega t - i\Phi) , \qquad (A1.32)$$

for the electric field of the diagnostic photons where

$$\Phi = \frac{\omega}{c} \int_0^z \tilde{n}(\omega, z') dz' = \frac{\omega z}{c} + \frac{\omega}{c} \int_0^z [\tilde{n}(\omega, z') - 1] dz'$$
(A1.33)

equation A1.30 becomes

$$\underline{E} = E_{o} \exp\left[i\omega\left(t - \frac{z}{c}\right) - A\overline{N_{i}L}W(-u + i\alpha)\right]$$
(A1.34)

where $\overline{N_iL}$ is the line-integrated-density of the resonant species of the neutral gas in a Maxwell-Boltzmann equilibrium. Finally, this means that

the measured fringe shift, $(\Delta s / s)$, of the diagnostic photons is given by the equation

$$\left(\frac{\Delta s}{s}\right) = \frac{A\overline{N_{i}L}}{2\pi} Im[W(-u+i\alpha)] \qquad (A1.35)$$

APPENDIX 2

CALCULATING ROOTS AND FRINGE SHIFTS

The following is an abbreviated version of the Mathematica program used to calculate the roots and determine the fringe shift for certain values of line-density and dye laser wavelength. The steps are few yet repeated many times with only changes in the input values, thus an abbreviation is given here.

First the roots of equation 2.71 are found by the following program:

h[eps_, dlambda_, nl_]=Exp[-(+518.047*dlambda+0.00314*I eps)^2] (1+Erf[I(+518.047*dlambda+0.00314*I-eps)]) -((-eps 2.56 10^13/nl) + I*0.5642)/(+518.047*dlambda+0.00314*I-eps) start = + 0.0 + 0.0*I dl = 0.003

a1 = FindRoot[h[gam, dl, 1. 10^{13}] == 0, {gam, start}, MaxIterations -> 100] itt=First[a1]/%

h[First[itt],0.004,1. 10^13]

a2 = FindRoot[h[gam, dl, 2. 10^13] == 0, {gam,First[itt]}, MaxIterations -> 100] itt=First[a2] /.% h[First[itt],dl,2. 10^13]

a3= (etc. ... for the desired values of line-density)

PutAppend[dl,a1,a2,a3, ..., "roots.h.output.ee2"]

!!roots.h.output.ee2 (this lists them out, so they can be cut and pasted into a seperate file. The output is alphabetic, not numeric, so it must be altered manually to obtain a numeric list of roots)

The function "h" is defined so that the function is equal to zero and a "start" value for the root search is given along with the dye laser wavelength, "dl", in nm. The root is determined using Newton's Method and each root is checked by inserting the root into the function. If the value of "h" using the newly found root is very small (e.g., $\leq 1 \times 10^{-7}$, value checked visually), then the root is valid. This process is repeated for different values of line-density, saved in a file using "PutAppend". The same format is used for other values of dye laser wavelength and then all of the roots are displayed using "!!".

For the solution to equation 2.61, the definition of "h" is changed

```
h[gamma_, dlambda_, nl_] = - gamma - nl*3.9056*10^-14 *

((+518.047*dlambda+0.00314*I-gamma) *

Exp[-(+518.047*dlambda+0.00314*I-gamma)^2] *

(1+Erf[I(+518.047*dlambda+0.00314*I-gamma)]) +

(-518.047*dlambda+0.00314*I+gamma) *

Exp[-(-518.047*dlambda+0.00314*I+gamma)^2] *

(1+Erf[I(-518.047*dlambda+0.00314*I+gamma)]))
```

to

The values {518.047}, {0.00314}, and { $3.9056x10^{-14}$ } in the above equations are for the Al(I), 394.401 nm, transition assuming a quasi-equilibrium, Maxwellian temperature of 0.3 eV. Note: $\Delta = (-518.047* \text{ dlambda})$ for [dlambda] = nm.

The roots are then reformated (alphabetic to numeric) and fed into the following program to determine fringe shift. The roots to equation 2.61 (gamma) are given here only for completeness.

(* if you change alpha, you also have to recalculate all the roots *) alpha = 0.00314 dlambda = 0.003 delta = -518.047*dlambda phi = +518.047*dlambda+alpha*I (* phi = -delta + I*alpha - gamma/epsilon *) chi = -518.047*dlambda+alpha*I (* chi = +delta + I*alpha + gamma/epsilon *) w[z_] = Exp[-(z^2)] * (1 + Erf[I*z]) mr[nl_, ue_] = Re[-5.9248 ue^2 - nl 2.314 10^-13 w[phi-ue]] iii[nl_, ue_] = Im[-5.9248 ue^2 - nl 2.314 10^-13 w[phi-ue]]

> (* the roots of eqn. 2.71 at a wavelength 0.003 nm. "e113" is the value of epsilon for the line-density 1x10¹³ cm⁻². *)

dlambda = 0.003
e113=-0.0488692 - 0.0547275*I
e213=-0.0899086 - 0.0991065*I
e313=-0.125766 - 0.136817*I
e413=-0.157883 - 0.169817*I
e513=-0.187145 - 0.199285*I
e613=-0.214138 - 0.225995*I
e713=-0.239277 - 0.250482*I
e813=-0.262864 - 0.273132*I
e913=-0.285131 - 0.294238*I
e114=-0.306256 - 0.314024*I
e214=-0.476653 - 0.466151*I
e314=-0.605138 - 0.573543*I
e414=-0.711273 - 0.65848*I
e514=-0.803064 - 0.729634*I
e614=-0.884692 - 0.791365*I
e714=-0.958649 - 0.846173*I
e814=-1.02659 - 0.895716*I
e914=-1.08962 - 0.941045*I
e115=-1.14858 - 0.982936*I
e215=-1.60344 - 1.29332*I
e315=-1.93096 - 1.50726*I
e415=-2.19515 - 1.67588*I
e515=-2.42009 - 1.8174*I
e615=-2.6177 - 1.94042*I
e715=-2.79535 - 2.05027*I
e815=-2.9573 - 2.14981*I
e915=-3.10665 - 2.24117*I

(* the roots of eqn. 2.61 at a wavelength 0.003 nm. "u113" is the value of gamma for the line-density 1×10^{13} cm⁻². *) dlambda = 0.003u113=-0.08496619504751728514 - 0.6031600065665911152*I u213=-0.3287336988152596157 - 0.9348385102125123111*I u313=-0.4157594146755698562 - 1.11150342490483498*I u413=-0.4287888815236202112 - 1.221998361485870857*I u513=-0.4191342770044838794 - 1.286263394630090693*I u613=-0.4067580400263541435 - 1.324476587177321576*I u713=-0.3961177066726843457 - 1.348931406109431985*I u813=-0.387566875961314686 - 1.365720590623060003*I u913=-0.3807266757810139003 - 1.37790624074367299*I u114=-0.3751930324789996254 - 1.387137341517644663*I u214=-0.3503887742179140818 - 1.423691049793081242*I u314=-0.3423466874362853795 - 1.434323734296974958*I u414=-0.3383895101318873075 - 1.4393925770574698*I u514=-0.3360372432247489145 - 1.442359651139959834*I u614=-0.3344785402432459137 - 1.444307854152442616*I u714=-0.3333698990835633334 - 1.445685143136329258*I u814=-0.3325410256628195829 - 1.446710421720261326*I u914=-0.3318979025210007285 - 1.447503358243030367*I

u115=-0.3313843894490563171 - 1.448134896706126034*I u215=-0.3290845085317007846 - 1.450946390829147029*I u315=-0.3283218879021801245 - 1.451872668165625699*I u415=-0.3279413331011181412 - 1.452333798993440078*I u515=-0.3277132427422407516 - 1.452609838765402512*I u615=-0.3275612836900032736 - 1.452793600074777329*I u715=-0.3274527911122523638 - 1.452924728549267829*I u815=-0.3273714488208655585 - 1.453023004140050014*I
(* Calculating Fringe Shift (FS) *)
fsa=iii[1. 10^{13} , $e113$] * 0.1592 fsb=iii[2. 10^{13} , $c213$] * 0.1592 fsc=iii[3. 10^{13} , $e313$] * 0.1592 fsc=iii[5. 10^{13} , $e413$] * 0.1592 fsf=iii[6. 10^{13} , $e613$] * 0.1592 fsf=iii[6. 10^{13} , $e613$] * 0.1592 fsh=iii[8. 10^{13} , $e813$] * 0.1592 fsh=iii[9. 10^{13} , $e813$] * 0.1592 fsi=iii[9. 10^{13} , $e813$] * 0.1592 fsk=iii[2. 10^{14} , $e214$] * 0.1592 fsk=iii[2. 10^{14} , $e214$] * 0.1592 fsk=iii[3. 10^{14} , $e314$] * 0.1592 fsn=iii[4. 10^{14} , $e414$] * 0.1592 fsn=iii[5. 10^{14} , $e514$] * 0.1592 fsn=iii[6. 10^{14} , $e614$] * 0.1592 fsn=iii[6. 10^{14} , $e614$] * 0.1592 fsn=iii[7. 10^{14} , $e714$] * 0.1592 fsn=iii[8. 10^{14} , $e814$] * 0.1592 fsn=iii[9. 10^{14} , $e914$] * 0.1592 fsn=iii[1. 10^{15} , $e115$] * 0.1592 fst=iii[2. 10^{15} , $e215$] * 0.1592 fst=iii[2. 10^{15} , $e315$] * 0.1592 fsv=iii[3. 10^{15} , $e415$] * 0.1592 fsv=iii[4. 10^{15} , $e415$] * 0.1592 fsv=iii[6. 10^{14} , $e514$] * 0.1592 fsv=iii[6. 10^{15} , $e615$] * 0.1592
fszaa=iii[8. 10^15, e815] * 0.1592 fszbb=iii[9. 10^15, e915] * 0.1592
PutAppend[dlambda, aaee,e113,e213,e313,e413,e513,e613,e713,e813,e913, e114,e214,e314,e414,e514,e614,e714,e814,e914, e115,e215,e315,e415,e515,e615,e715,e815,e915, aauu,u113,u213,u313,u413,u513,u613,u713,u813,u913, u114,u214,u314,u414,u514,u614,u714,u814,u914, u115,u215,u315,u415,u515,u615,u715,u815,u915, aafs,fsa,fsb,fsc,fsd,fse,fsf,fsg,fsh,fsi,fsj,fsk,fsl,fsm,fsn,fso,fsp,fsq,fsr, fss,fst,fsu,fsv,fsx,fsy,fsz,fszaa,fszbb, "roots.h.output.f003"]

!!roots.h.output.f003

- (* This is the Sept. 1993 streamline model of the calculation of the fringe shift. Just enter the temperature and the wavelength in at the top, and the program will give you the fringe shift for the preselected line-densities. Next, linearly interpolate to find the line-density for a given fringe shift. Roger Alan Lindley *)
- (* the notation "e113" is to be read as the epsilon root for the line density 1x10^13 cm-2 *)

temp=3.0 (* enter in the temperature in [eV] *) dl=0.006 (* enter in the wavelength in [nm] *)

alpha=0.00314*(0.3/temp)^0.5 ddd=518.047*(0.3/temp)^0.5 kkkk=2.56 10^13*(temp/0.3)^1.5 cccc=5.9248*(temp/0.3)^1.0 aaaa=2.314 10^-13*(0.3/temp)^0.5

 $w[z_] = Exp[-(z^2)] * (1 + Erf[I*z])$

 $rrr[nl_, ue_] = Re[-cccc^*ue^2 - nl^*aaaa * w[dddd^*dl+alpha*I-ue]]$ iii[nl_, ue_] = Im[-cccc^*ue^2 - nl^*aaaa * w[dddd^*dl+alpha*I-ue]]

(* MEASURES' ANSWER for line-density per fringe shift: *) nldss=(2*3.14159/aaaa)/(Im[w[+dddd*dl+alpha*I]])

h[eps_, dl_, nl_]=w[+dddd*dl+alpha*I-eps] - ((-eps*kkkk/nl) + I*0.5642)/(+dddd*dl+alpha*I-eps)

a1 = FindRoot[h[eps, dl, 1. 10^13] == 0, {eps,0.0}, MaxIterations -> 100] itt=First[a1] /.% e113=First[itt] h[e113,dl,1. 10^13]

a2 = FindRoot[h[eps, dl, 2. 10^13] == 0, {eps,e113}, MaxIterations -> 100] itt=First[a2] /.% e213=First[itt] h[e213,dl,2. 10^13]

a3 = FindRoot[h[eps, dl, 3. 10^13] == 0, {eps,e213}, MaxIterations -> 100] itt=First[a3] /.% e313=First[itt] h[e313,dl,3. 10^13]

(* the program is abbreviated here, continue the above for all of the desires values of line density *)

 $fsa=iii[1. 10^{13}, e113] * 0.1592$ $fsb=iii[2. 10^{13}, e213] * 0.1592$ $fsc=iii[3. 10^{13}, e313] * 0.1592$ $fsd=iii[4. 10^{13}, e413] * 0.1592$

fse=iii[5. 10^13, e513] * 0.1592 fsf=iii[6. 10^13, e613] * 0.1592 fsg=iii[7. 10^13, e713] * 0.1592 fsh=iii[8. 10^13, e813] * 0.1592 fsi=iii[9. 10^13, c913] * 0.1592 fsj=iii[1. 10^14, c114] * 0.1592 fsk=iii[2. 10^14, e214] * 0.1592 $fsl=iii[3, 10^{14}, e314] * 0.1592$ fsm=iii[4. 10^14, e414] * 0.1592 fsn=iii[5. 10^14, e514] * 0.1592 fso=iii[6. 10^14, e614] * 0.1592 fsp=iii[7. 10^14, e714] * 0.1592 fsq=iii[8. 10^14, e814] * 0.1592 fsr=iii[9. 10^14, e914] * 0.1592 fss=iii[1. 10^15, e115] * 0.1592 fst=iii[2. 10^15, e215] * 0.1592 fsu=iii[3. 10^15, e315] * 0.1592 fsv=iii[4. 10^15, e415] * 0.1592 fsx=iii[5. 10^15, e515] * 0.1592 fsy=iii[6. 10^15, e615] * 0.1592 fsz=iii[7. 10^15, e715] * 0.1592 fszaa=iii[8. 10^15, e815] * 0.1592 fszbb=iii[9. 10^15, e915] * 0.1592

PutAppend[dl,temp, aaee,e113,e213,e313,e413,e513,e613,e713,e813,e913,e114,e214,e314,e414,e514,e61 4,e714,e814,e914,e115,e215,e315,e415,e515,e615,e715,e815,e915, aafs,fsa,fsb,fsc,fsd,fse,fsf,fsg,fsh,fsi,fsj,fsk,fsl,fsm,fsn,fso,fsp,fsq,fsr,fss,fst,fsu,fsv ,fsx,fsy,fsz,fszaa,fszbb, "roots.h.output.final.1"]

!!roots.h.output.final.1

APPENDIX 3 FRINGE SHIFT INTERPRETATION GRAPHS

The following graphs represent the fringe shift interpretations for the holograms presented in Chapter 4 as calculated by Griem (eqn. 2.9) assuming an infinitesimal laser bandwidth and no Doppler shift effects (cold ablation plume), by Measures (eqn. 2.86) accounting for the Doppler shift effects and assuming an infinitesimal laser bandwidth, by Dreiden (eqn. 2.28) accounting for a finite laser bandwidth but assuming no Doppler shift effects and no selective absorption effects, and by this work (eqn. 2.85) accounting for Doppler shift effects, finite laser bandwidth, and selective absorption effects.



Fig. A3.1 Fringe shift interpretation at 0.003 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.2 Fringe shift interpretation at 0.004 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.3 Fringe shift interpretation at 0.006 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.4 Fringe shift interpretation at 0.008 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.5 Fringe shift interpretation at 0.009 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.6 Fringe shift interpretation at 0.013 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.7 Fringe shift interpretation at 0.015 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.8 Fringe shift interpretation at 0.016 nm for the Al(I) 394.401 nm line assuming a plume temperature of 0.3 eV.



Fig. A3.9 Measures' calculation at 0.003 nm for several ablation plume temperatures.



Fig. A3.10 This work's calculation at 0.003 nm for several ablation plume temperatures.



Fig. A3.11 Measures' calculation at 0.006 nm for several ablation temperatures.



Fig. A3.12 This work's calculation at 0.006 nm for several ablation plume temperatures.



Fig. A3.13 Measures' calculation at 0.009 nm for several ablation plume temperatures.



Fig. A3.14 This work's calculation at 0.009 nm for several ablation plume temperatures.

APPENDIX 4 TIME DELAY CIRCUITS

The time delay between the Lumonics Laser and the Lambda Physik Excimer pumped Dye Laser was varied using one of two electronic arrangements. Figure A4.1 shows the timing circuit used for the holographic data presented in the text from October 1992 through April 1993, and Figure A4.2 shows the timing circuit used for the holographic data presented in the text for May through June 1993. The only difference is the Delay Generator setup (shown at the top of the figures).

The time delay between the two laser pulses was measured on the Tektronix DSA 602. Figure A4.3 shows an example of the recorded traces. For examples of the calorimeter signals, see Appendix 5.



Fig. A4.1 Electrical schematic for July 1992 to April 1993.



Fig. A4.2 Electrical schematic for May to June, 1993.



Fig. A4.3 Example of the timing signals as seen from the Digital Signal Analyzer (DSA) on a scale of 500 ns/div. The upper trace (2 mV/div) is the PIN diode signal. The isolated peak on the right side is the dye laser signal, while the noise at earlier times is due to the electromagnetic emission of the KrF laser internal electrical discharge. The lower trace (500 mV/div) is the silicon photo-diode signal. The initial edge of the of the signal represents the KrF laser pulse, while the long tail is either due to the fluorescence of the KrF laser beam focusing lens or to the supporting circuitry of the silicon photo-diode (RC decay). The time difference is determined by aligning a vertical bar with each of the laser peaks (here shown as two diamonds on the lower trace) and reading the DSA time difference below (here, 1.350 μ s).

APPENDIX 5 ENERGY CALIBRATION

The shot-to-shot energy is monitored by splitting off part of the Lumonics laser beam onto a calorimeter, called the Side Calorimeter (Gen-Tec LED-200-C, 8.8 V/J, diameter ≈ 2 cm). It is general practice to calibrate the signal from this calorimeter after each experimental run because: 1) the black, front surface of the Side Calorimeter has been irregularly ablated over its lifetime of use, and 2) because moving the focal lens changes the absorption spot on the Side Calorimeter as well as the ablation spot on the target. Figure 3.1(b) shows the experimental configuration for this calibration, where the Side Calorimeter is calibrated using a second calorimeter (Lumonics 50D-171, 0.39 V/J, diameter ≈ 11.5 cm).

A Tektronix 5223 Digitizing Oscilloscope was used to record the calorimeter signals. Figure A5.1 shows an examples of the calorimeter traces during a calibration run. The peak signals are divided by the appropriate conversion factors (given above) to obtain the incident energy. Figure A5.2 shows an example calibration graph, where the side calorimeter energy is graphed versus the energy hitting the target (back calorimeter) (see Fig. 3.1(b)).

The same target spot on the same aluminum sample was used during the experimental runs during October 3 - 23, 1992 (see Fig. 3.2). Unfortunately, a side-to-back calibration was not performed after the last run and before the optical table components were disturbed. The energy calibration for the data taken in that time period was determined in the following way. It is a general practice to place the larger calorimeter in



Fig. A5.1 Example of calorimetry signals as seen from the Tektronix 5223 Digitizing Oscilloscope on a scale of 20 ms/div. The top trace (50 mV/div) is the side calorimeter at position (C) on Fig. 3.1(b) and the bottom trace (50 mV/div) is the second calorimeter placed at position (D). Note: the lower trace has the higher peak and undershoots after \approx 55 ms.



Fig. A5.2 Example fluence calibration of the side calorimeter (position C) versus the energy measured behind the chamber (position D): deposition in vacuum experiment, June 17, 1993; (see Fig. 3.1(b))



Fig. A5.3 Calibration of the energy measured behind the chamber (position D) to calibrated to the front of the chamber (position F); (see Fig. 3.1(b)).

front of the focusing lens for a single shot both before and after taking a holographic plate for comparison with the total Lumonics laser output since the laser's purchase and installation. The larger calorimeter in this position is called the Front Calorimeter. In an experiment in July 1992, both Front and Side calorimetry signals were recorded while taking the holographic plates and a Side-to-Back calibration was then performed. Figure A5.4 shows the Front-to-Back calibration curve determined from this July 1992 experiment and used with the October 1992 data as an approximation of the energy hitting the target.

APPENDIX 6

ANALYZING THE LASER ABLATION PLUME EXPANSION

Figures A6.1 through A6.5 show the plume expansion analysis, as described in Chapter 4, for the holograms in Figures 4.6, 4.8, 4.12, 4.14, and 4.16. Each connected set of points represents one hologram, and each point represents the farthest extension of an equicontour line for the hologram. For the information derived from these graphs, see Figures 4.20, 4.21, 4.22, and 4.23.



Fig. A6.2 Analysis of plume expansion in vacuum for the deposition case; for the holograms in Fig. 4.8.



Fig. A6.4 Analysis of plume expansion in 52 mTorr argon gas; for the holograms in Fig. 4.14.



Fig. A6.5 Analysis of plume expansion in 210 mTorr argon gas; for the holograms in Fig. 4.16.

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